

Naturalness, Dark Matter and E_6 Inspired SUSY Models

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Based on:

P. Athron, D. Harries, and A. G. Williams, Phys. Rev. **D91**, 115024 [arXiv:1503.08929]

P. Athron, D. Harries, R. Nevzorov, and A. G. Williams, Phys. Lett. **B760**, 19 (2016) [arXiv:1512.07040]

P. Athron, D. Harries, R. Nevzorov, and A. G. Williams, JHEP **12**, 128 (2016) [arXiv:1610.03374]

P. Athron, C. Balázs, B. Farmer, A. Fowlie, D. Harries, and D. Kim, arXiv:1709.07895

November 2, 2017

Outline

Motivation

E_6 Inspired Models

Fine Tuning in E_6 Models

The SE_6SSM

Dark Matter in the CSE_6SSM

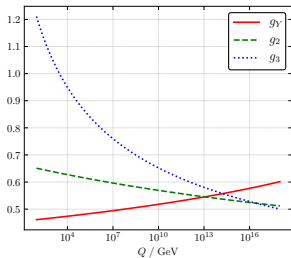
Summary

The Case for BSM Physics



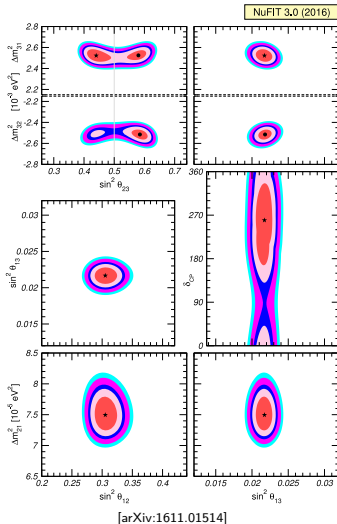
[APOD/NASA]

Dark matter?

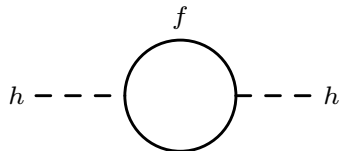


[<http://flexiblesusy.hepforge.org/images.html>]

Gauge unification?



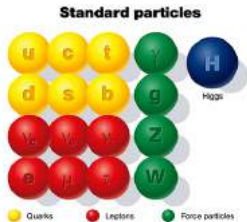
Neutrino masses?



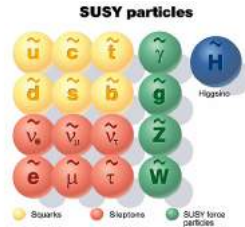
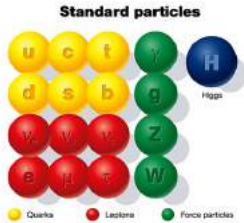
$$m_h^2 \ll \Lambda_{NP}^2?$$

i.e., the "Hierarchy Problem"?

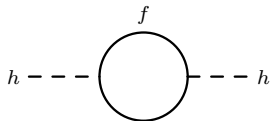
The MSSM: A Schematic View



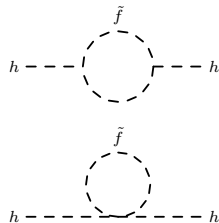
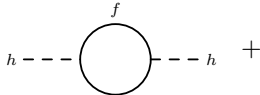
(almost)



[<http://www.physics.gla.ac.uk/ppt/bsm.htm>]



→



MSSM Field Content

- ▶ **Minimal** extension of the SM compatible with SUSY
- ▶ SM matter, gauge fields \Rightarrow supermultiplets
- ▶ Additional Higgs doublet required
- ▶ Characterised by superpotential + soft SUSY breaking interactions
 - ▶ in general \Rightarrow **large number of new parameters**

	$s = 0$	$s = \frac{1}{2}$	$SU(3)_C$	$SU(2)_L$	$\sqrt{\frac{5}{3}} Q_i^Y$
\hat{Q}_i	$\begin{pmatrix} \tilde{u}_L \\ \tilde{d}_L \end{pmatrix}_i$	$\begin{pmatrix} u_L \\ d_L \end{pmatrix}_i$	3	2	$\frac{1}{6}$
\hat{u}_i^c	\tilde{u}_{iR}^*	u_{iR}^c	$\bar{\mathbf{3}}$	1	$-\frac{2}{3}$
\hat{d}_i^c	\tilde{d}_{iR}^*	d_{iR}^c	$\bar{\mathbf{3}}$	1	$\frac{1}{3}$
\hat{L}_i	$\begin{pmatrix} \tilde{\nu}_L \\ \tilde{e}_L \end{pmatrix}_i$	$\begin{pmatrix} \nu_L \\ e_L \end{pmatrix}_i$	1	2	$-\frac{1}{2}$
\hat{e}_i^c	\tilde{e}_{iR}^*	e_{iR}^c	1	1	1
\hat{H}_d	$\begin{pmatrix} H_d^0 \\ H_d^- \end{pmatrix}$	$\begin{pmatrix} \tilde{H}_d^0 \\ \tilde{H}_d^- \end{pmatrix}$	1	2	$-\frac{1}{2}$
\hat{H}_u	$\begin{pmatrix} H_u^+ \\ H_u^0 \end{pmatrix}$	$\begin{pmatrix} \tilde{H}_u^+ \\ \tilde{H}_u^0 \end{pmatrix}$	1	2	$\frac{1}{2}$

$$W_{\text{MSSM}}^{\text{RPC}} = \mu(\hat{H}_d \cdot \hat{H}_u) + y_{ij}^e \hat{e}_i^c (\hat{L}_j \cdot \hat{H}_d) + y_{ij}^d \hat{d}_i^c (\hat{Q}_j \cdot \hat{H}_d) + y_{ij}^u \hat{u}_i^c (\hat{H}_u \cdot \hat{Q}_j)$$

DM in the (RPC) MSSM

- ▶ Most general superpotential $\supset B, L$ violating interactions
- ▶ Forbid most dangerous terms? \Rightarrow impose matter parity ($\equiv R$ -parity):

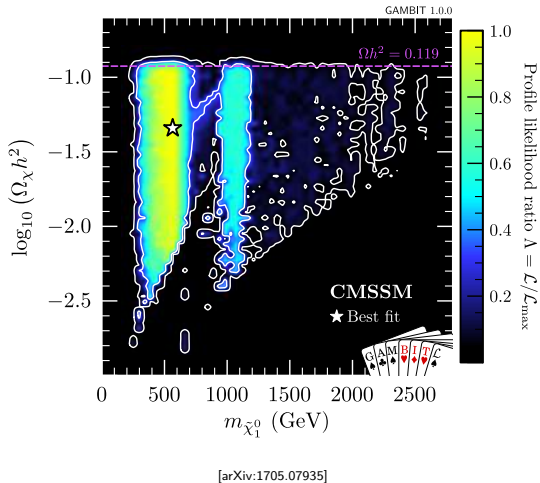
$$Z_2^M = (-1)^{3(B-L)}$$

- ▶ R -parity conservation \Rightarrow electrically neutral LSP is stable, natural WIMP DM candidate

- ▶ Standard MSSM candidate is lightest $\tilde{\chi}^0$,

$$\tilde{\chi}_1^0 = N_{11}\tilde{H}_d^0 + N_{12}\tilde{H}_u^0 + N_{13}\tilde{W}_3 + N_{14}\tilde{B}$$

- ▶ Mixings N_{1i} \Rightarrow different DM scenarios (“bino-like”, “wino-like”, ...)





Combined Measurement of the Higgs Boson Mass in pp Collisions at $\sqrt{s} = 7$ and 8 TeV with the ATLAS and CMS Experiments

G. Aad *et al.**

(ATLAS Collaboration)[†]

(CMS Collaboration)[‡]

(Received 25 March 2015; published 14 May 2015)

A measurement of the Higgs boson mass is presented based on the combined data samples of the ATLAS and CMS experiments at the CERN LHC in the $H \rightarrow \gamma\gamma$ and $H \rightarrow ZZ \rightarrow 4\ell$ decay channels. The results are obtained from a simultaneous fit to the reconstructed invariant mass peaks in the two channels and for the two experiments. The measured masses from the individual channels and the two experiments are found to be consistent among themselves. The combined measured mass of the Higgs boson is $m_H = 125.09 \pm 0.21$ (stat) ± 0.11 (syst) GeV.

DOI: 10.1103/PhysRevLett.114.191803

PACS numbers: 14.80.Bn, 13.85.Qk

MSSM tree-level prediction:

$$m_{h_1}^2 \leq m_Z^2 \cos^2 2\beta \lesssim (91 \text{ GeV})^2$$

The "Little Hierarchy Problem"

- ▶ $m_{h_1} \approx 125 \text{ GeV} \Rightarrow$ large higher order corrections

$$m_{h_1}^2 \approx m_Z^2 \cos^2 2\beta \left(1 - \frac{3}{8\pi^2} \frac{m_t^2}{v^2} \ln \frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{M_t^2} \right) + \frac{3}{4\pi^2} \frac{m_t^4}{v^2} \ln \frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{M_t^2} + \dots$$

- ▶ \Rightarrow also large corrections to prediction for m_Z at SUSY scale M_S :

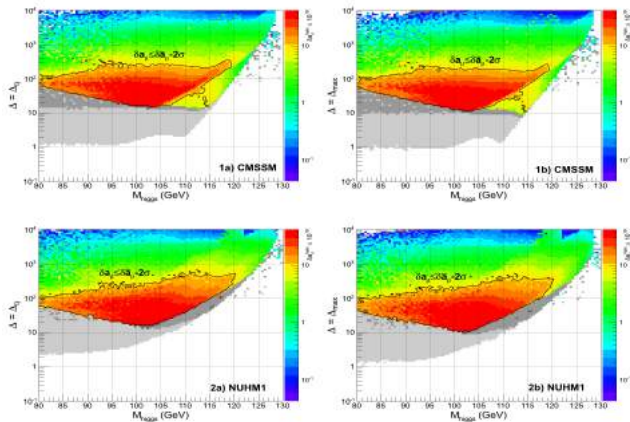
$$\frac{m_Z^2}{2} = -\mu^2 + \overbrace{\frac{m_{H_d}^2 - m_{H_u}^2}{\tan^2 \beta - 1} \tan^2 \beta}^{\text{RGE effects}} + \delta_{1\text{-loop}},$$

$$\delta_{1\text{-loop}} = \frac{3}{8\pi^2} \frac{m_t^2}{v^2 \cos 2\beta} \left[m_{\tilde{t}_1}^2 \left(\ln \frac{m_{\tilde{t}_1}^2}{M_S^2} - 1 \right) + m_{\tilde{t}_2}^2 \left(\ln \frac{m_{\tilde{t}_2}^2}{M_S^2} - 1 \right) \right] + \dots$$

- ▶ \Rightarrow naturalness problem?
- ▶ **μ -problem:** $m_Z^2/2 = -\mu^2 + \dots \Rightarrow \mu \sim$ soft parameters?

The "Little Hierarchy Problem"

- ▶ E.g., constrained models \Rightarrow large *sensitivities* Δ



[arXiv:1203.0569]

Non-minimal SUSY Models

- ▶ Motivated by MSSM shortcomings such as LHP, μ -problem
 - ▶ also, e.g., non-zero ν masses, baryogenesis, ...
- ▶ Extend MSSM matter content and/or gauge symmetries
- ▶ Simplest example: NMSSM = MSSM + SM singlet superfield,

$$W_{\text{NMSSM}}^{Z_3} = W_{\text{MSSM}}^{\text{RPC}}|_{\mu=0} + \lambda \hat{S} \hat{H}_d \cdot \hat{H}_u + \frac{1}{3} \kappa \hat{S}^3$$

$$m_{h_1}^2 \leq m_Z^2 \cos^2 2\beta + \frac{\lambda^2 v^2}{2} \sin^2 2\beta$$

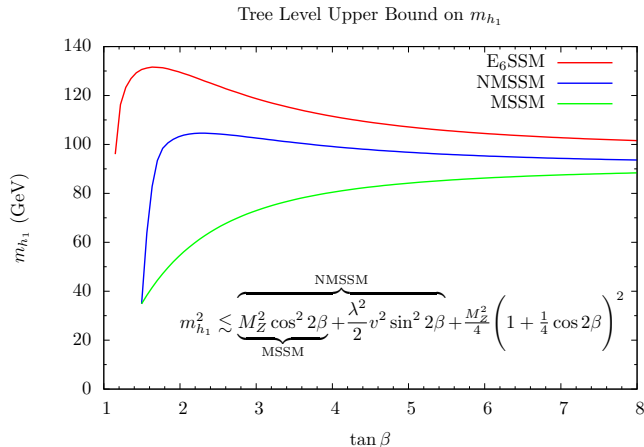
$$\langle S \rangle \neq 0 \Rightarrow \mu_{\text{eff.}} = \lambda \langle S \rangle$$

- ▶ Not without complications, may suffer from additional problems
 - ▶ e.g., domain walls in the NMSSM

$U(1)$ Extensions of the MSSM

$$SU(3)_C \times SU(2)_L \times U(1)_Y \rightarrow SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)'$$

- ▶ Well-motivated extensions of MSSM, NMSSM
- ▶ $W_{\text{USSM}} \supset \lambda \hat{S} \hat{H}_d \cdot \hat{H}_u, Q'_S \neq 0 \Rightarrow \langle S \rangle = s/\sqrt{2}$ also breaks $U(1)'$, generating **massive Z'**
- ▶ Consistent model requires anomaly cancellation
 - ▶ either family non-universal $U(1)'$ charges or **extra matter**
- ▶ Additional states \Rightarrow exciting phenomenology



E_6 Inspired Models

- ▶ Lead to $U(1)$ extended models at low-energies:

$$\begin{aligned} E_6 &\longrightarrow SO(10) \times U(1)_\psi \\ &\longrightarrow SU(5) \times U(1)_\psi \times U(1)_\chi \\ &\longrightarrow SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_\psi \times U(1)_\chi \\ &\longrightarrow SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)' \end{aligned}$$

- ▶ Resulting charges $Q' = Q_\chi \cos \theta_{E_6} + Q_\psi \sin \theta_{E_6}$, e.g., class of models

$$U(1)_N: \quad Q_N = Q(\theta_{E_6} = \arctan \sqrt{15}) \quad (\equiv E_6\text{SSM})$$

$$U(1)_\psi: \quad Q_\psi = Q(\theta_{E_6} = \pi/2)$$

$$U(1)_\eta: \quad Q_\eta = -Q(\theta_{E_6} = \pi - \arctan \sqrt{5/3})$$

$$U(1)_I: \quad Q_I = -Q(\theta_{E_6} = \arctan \sqrt{3/5})$$

- ▶ Matter content fills complete **27** representations (**ensures anomaly cancellation**)
 - ▶ \Rightarrow additional exotic states

The E_6 SSM

- ▶ $\tan \theta_{E_6} = \sqrt{15} \Rightarrow U(1)_N$ under which right-handed neutrinos are uncharged
 - ▶ allows ν masses via see-saw and successful baryogenesis [1]
- ▶ Extra $\hat{L}_4, \hat{\bar{L}}_4$ from incomplete $\mathbf{27}'$, $\overline{\mathbf{27}}'$ for gauge unification
- ▶ Low-energy matter content from $\mathbf{27}$ -plet:

$$(\hat{Q}_i, \hat{u}_i^c, \hat{d}_i^c, \hat{L}_i, \hat{e}_i^c) + (\hat{D}_i, \hat{\bar{D}}_i) \\ + (\hat{S}_i) + (\hat{H}_i^u) + (\hat{H}_i^d)$$

- ▶ Higgs doublets \hat{H}_3^d, \hat{H}_3^u and one singlet \hat{S}_3 get VEVs (\Rightarrow EWSB and break $U(1)_N$)

$$W_{E_6\text{SSM}} \approx y_\tau \hat{L}_3 \cdot \hat{H}_3^d \hat{e}_3^c + y_b \hat{Q}_3 \cdot \hat{H}_3^d \hat{d}_3^c + y_t \hat{H}_3^u \cdot \hat{Q}_3 \hat{u}_3^c + \lambda_i \hat{S}_3 \hat{H}_i^d \cdot \hat{H}_i^u + \kappa_i \hat{S}_3 \hat{D}_i \hat{\bar{D}}_i + \mu_L \hat{L}_4 \cdot \hat{\bar{L}}_4$$

	$SU(3)_C$	$SU(2)_L$	$\sqrt{\frac{5}{3}} Q_i^Y$	$\sqrt{40} Q_i^N$
\hat{Q}_i	$\mathbf{3}$	$\mathbf{2}$	$\frac{1}{6}$	1
\hat{u}_i^c	$\overline{\mathbf{3}}$	$\mathbf{1}$	$-\frac{2}{3}$	1
\hat{d}_i^c	$\overline{\mathbf{3}}$	$\mathbf{1}$	$\frac{1}{3}$	2
\hat{L}_i	$\mathbf{1}$	$\mathbf{2}$	$-\frac{1}{2}$	2
\hat{e}_i^c	$\mathbf{1}$	$\mathbf{1}$	1	1
\hat{S}_i	$\mathbf{1}$	$\mathbf{1}$	0	5
\hat{H}_i^u	$\mathbf{1}$	$\mathbf{2}$	$\frac{1}{2}$	-2
\hat{H}_i^d	$\mathbf{1}$	$\mathbf{2}$	$-\frac{1}{2}$	-3
\hat{D}	$\mathbf{3}$	$\mathbf{1}$	$-\frac{1}{3}$	-2
$\hat{\bar{D}}$	$\overline{\mathbf{3}}$	$\mathbf{1}$	$\frac{1}{3}$	-3
\hat{L}_4	$\mathbf{1}$	$\mathbf{2}$	$-\frac{1}{2}$	2
$\hat{\bar{L}}_4$	$\mathbf{1}$	$\overline{\mathbf{2}}$	$\frac{1}{2}$	-2

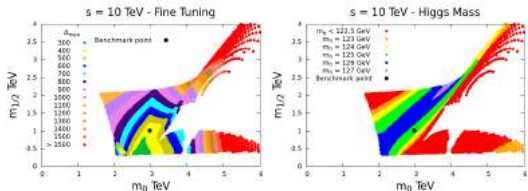
The Impact of D -terms

$$m_{h_1}^2 \leq m_Z^2 \cos^2 2\beta + \frac{\lambda^2 v^2}{2} \sin^2 2\beta + \frac{m_Z^2}{4} \left(1 + \frac{1}{4} \cos 2\beta\right)^2$$

$$c \frac{m_Z^2}{2} \approx -\mu_{\text{eff}}^2 + \frac{\bar{m}_{H_d}^2 - \bar{m}_{H_u}^2 \tan^2 \beta}{\tan^2 \beta - 1} + d \frac{m_{Z'}^2}{2}$$

- ▶ Extra D -terms \Rightarrow reduced radiative corrections
- ▶ E.g., CE₆SSM shows reduced sensitivity compared to CMSSM

- ▶ Large $m_{Z'}$ \Rightarrow large contribution to m_Z
- ▶ $m_{Z'} \gg m_Z \Rightarrow$ fine-tuning again?
- ▶ Coefficient $d = d(\tan \beta; \theta_{E_6}) \Rightarrow$ impact of $m_{Z'}$ depends on θ_{E_6}
 - ▶ $c = c(\tan \beta; \theta_{E_6}) = O(1)$



[arXiv:1302.5291]

For $m_{Z'}$ large, does cost outweigh the benefit?

Model Dependent Contributions

- ▶ Radiative contributions, e.g.,

$$m_{H_u}^2(M_S) \sim m_{H_u}^2(M_X) + \frac{3y_t^2}{8\pi^2} [m_{H_u}^2(M_X) + m_{Q_3}^2(M_X) + A_t^2(M_X)] \ln \frac{M_S}{M_X} + \dots$$

depend on

- ▶ soft SUSY breaking mechanism (\Leftrightarrow values of $m_{H_u}^2(M_X)$, $m_{Q_3}^2(M_X)$, etc.)
- ▶ scale of SUSY breaking, M_X
- ▶ I.e., conclusions strongly depend on assumptions about (unknown) high-energy boundary conditions
- ▶ Contrast with *tree-level* $m_{Z'}$ effect
- ▶ Study models at low energies \Rightarrow conservative picture of tuning
- ▶ Alternative: measure fine-tuning purely in terms of low-energy parameters?

Measuring Fine-tuning?

What constitutes "fine-tuning"?
Opinions differ.

- ▶ Large cancellations?

$$\Delta_{EW} = \max_i \frac{2|C_i|}{m_Z^2}, \quad C_1 = -\mu^2, \quad C_2 = \frac{m_{H_d}^2}{\tan^2 \beta - 1}, \dots$$

- ▶ Extreme sensitivities?

$$\Delta_{BG} = \max_i \left| \frac{\partial \ln m_Z^2}{\partial \ln p_i} \right|, \quad p_i \in \{\text{fundamental parameters}\}$$

- ▶ High or low energy contributions (Δ_{EW} or Δ_{HS})? Definition of parameters p_i ? Just m_Z ?
- ▶ Model fine-tuned $\Leftrightarrow \Delta_{EW/HS/BG/\dots} > ?$
- ▶ Compare tuning between models?

A Bayesian Approach

- ▶ Rigorous framework automatically captures intuition about "naturalness"
- ▶ Bayes' theorem applied to model M , parameters \mathbf{x} , "observables" \mathbf{O} :

$$p(\mathbf{x}|\text{data}, M) = \frac{p(\text{data}|\mathbf{x}, M)p(\mathbf{x}|M)}{p(\text{data}|M)}, \quad p(M|\text{data}) = \frac{p(\text{data}|M)p(M)}{p(\text{data})}$$

$$p(\text{data}|M) = \int d^n x_i p(\text{data}|\mathbf{x}, M)p(\mathbf{x}|M), \quad p_{\text{eff.}}(x_j, \dots) = \int dO_i p(\mathbf{O}, \mathbf{x}'|M, O_i^{\text{exp.}})$$

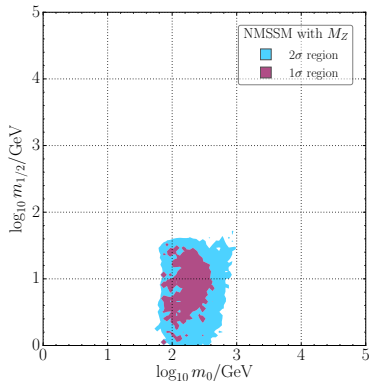
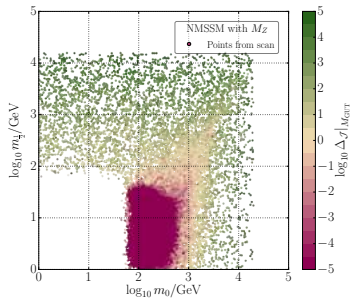
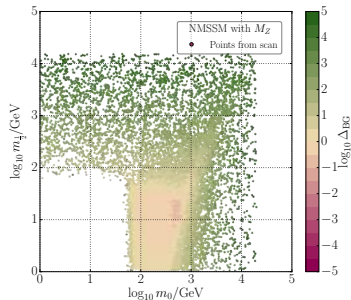
- ▶ Reparameterise in favour of \mathbf{O} , e.g.,

$$p(\mathbf{x}|M) = \mathcal{J}p(\mathbf{O}, \mathbf{x}'|M)$$

⇒ evidence and effective priors suppressed by Jacobian $\Delta_J \sim \mathcal{J}$,

$$\Delta_J = \left| \det \frac{\partial \ln O_i}{\partial \ln x_j} \right|$$

Example: Bayesian Naturalness in the NMSSM



[arXiv:1709.07895]

- ▶ Δ_{BG} , Δ_J qualitatively similar
- ▶ High posterior density \Leftrightarrow low fine-tuning
- ▶ Before $m_{h_1} \approx 125$ GeV, weak-scale soft parameters preferred

Fine Tuning in the E_6 SSM

- ▶ Use Δ_{BG} for simplicity and comparison with previous results
 - ▶ Expect similar conclusions to full Bayesian approach
- ▶ Fundamental model parameters \Leftrightarrow boundary values at M_X
 - ▶ e.g. $m_{H_u}^2(M_X)$, $A_t(M_X)$, $A_\lambda(M_X)$, $M_3(M_X)$, ...
- ▶ Minimise impact of high-energy assumptions \Rightarrow choose low $M_X = 20$ TeV
- ▶ Account for remaining impact with approximate solutions ($t = \ln \frac{M_S}{M_X}$, $\frac{dp}{dt} = \frac{\beta_p^{(1)}}{(4\pi)^2} + \frac{\beta_p^{(2)}}{(4\pi)^4}$)

$$p(M_S) \approx p(M_X) + \frac{t}{(4\pi)^2} \left(\beta_p^{(1)}(M_X) + \frac{1}{(4\pi)^2} \beta_p^{(2)}(M_X) \right) + \frac{t^2}{2(4\pi)^4} \sum_{\{q\}} \beta_q^{(1)}(M_X) \left. \frac{\partial \beta_p^{(1)}}{\partial q} \right|_{M_X}$$

- ▶ Dependence of Δ_{BG} on $m_{Z'}$?

Numerical Methods

- ▶ Consider multiple θ_{E_6} values (role of d coefficient?)
- ▶ "Phenomenological" boundary conditions at $M_X \Rightarrow$ compute tunings for

$$p = \{\lambda, A_\lambda, m_{H_d}^2, m_{H_u}^2, m_S^2, m_{Q_3}^2, m_{u_3^c}^2, A_t, \\ M_1, M_2, M_3, M_1'\}$$

- ▶ 2-loop RGEs generated by SARAH + FlexibleSUSY
- ▶ For MSSM used modified SOFTSUSY 3.3.10
 - ▶ Assume pMSSM boundary conditions with $\mu, B \in [-1, 1]$ TeV
- ▶ Neglect kinetic mixing
 - ▶ small in models considered
 - ▶ not negligible in general (e.g. $U(1)_{B-L}$)

$$2 \leq \tan \beta \leq 50 \text{ for } U(1)_N, \\ \tan \beta = 10 \text{ for all others}$$

$$-3 \leq \lambda \leq 3$$

$$-10 \text{ TeV} \leq A_\lambda \leq 10 \text{ TeV}$$

$$200 \text{ GeV} \leq m_{Q_3} \leq 2000 \text{ GeV}$$

$$200 \text{ GeV} \leq m_{u_3} \leq 2000 \text{ GeV}$$

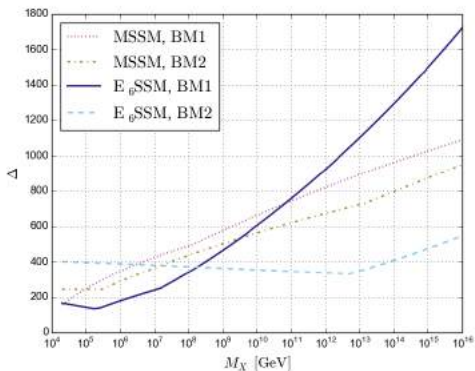
$$-10 \text{ TeV} \leq A_t \leq 10 \text{ TeV}$$

$$M_2 = \\ 100 \text{ GeV}, 1050 \text{ GeV}, 2000 \text{ GeV}$$

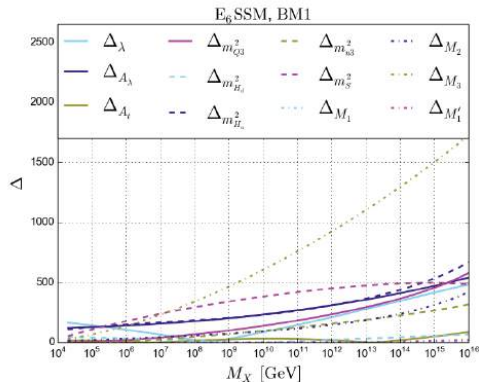
- ▶ Other soft scalar masses = 5 TeV, $M_1 = M_1' = 300$ GeV, $M_3 = 2$ TeV

High-scale Model Dependence

- ▶ Δ_{BG} highly model-dependent
 - ▶ E.g., E_6 SSM BM1 (unconstrained) vs. BM2 ($\sim CE_6$ SSM)

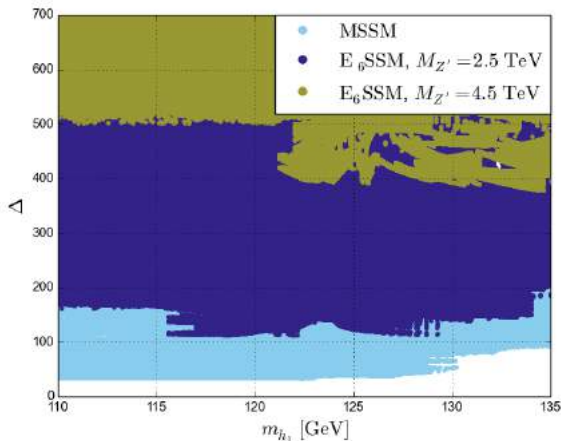


[arXiv:1503.08929]



- ▶ Large $M_X \Rightarrow$ RG contributions (e.g., M_3 in E_6 SSM BM1) are dominant
- ▶ High-scale models: impact of $m_{Z'}$ limits not easily seen

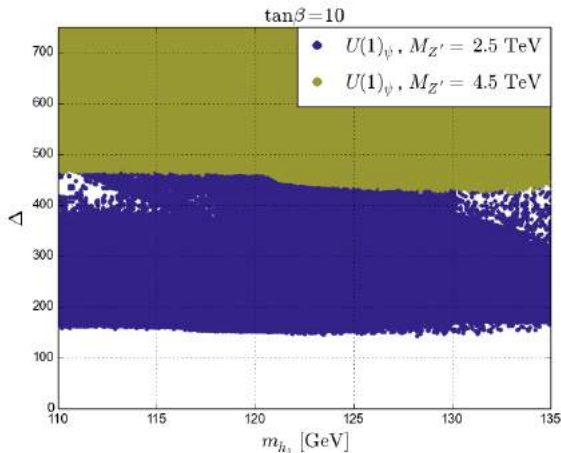
$U(1)_N$: $Q_N = Q(\theta_{E_6} = \arctan \sqrt{15})$, $d(\tan \beta = 10) \approx 0.40$



[arXiv:1503.08929]

- ▶ Low $M_X \Rightarrow \tilde{t}$ tuning reduced
- ▶ $m_{Z'}$ sets lower limit on Δ_{BG}
 - ▶ \Rightarrow conservative limit on naturalness in E_6SSM
- ▶ $m_{Z'} \approx 2.5$ TeV (old limit!) \Rightarrow minimum tuning already exceeds MSSM
 - ▶ \approx equivalent to MSSM with ≥ 700 GeV $\tilde{\chi}^\pm$ (\Rightarrow lower bound on μ)
- ▶ Warning: only 1-loop t, \tilde{t} + leading 2-loop Higgs mass calculation \Rightarrow large Higgs mass uncertainty

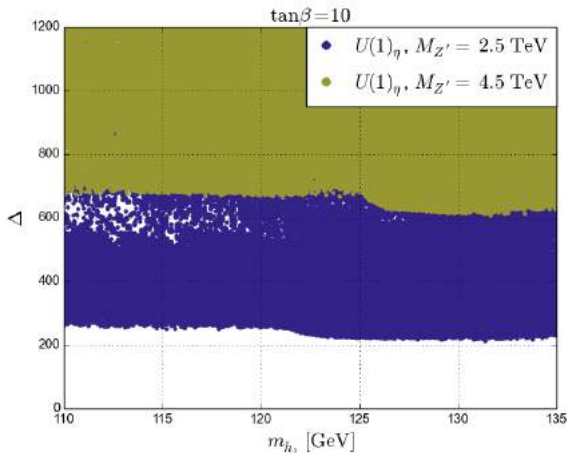
$$U(1)_\psi: Q_\psi = Q(\theta_{E_6} = \pi/2), d(\tan \beta = 10) \approx 0.50$$



[arXiv:1503.08929]

- ▶ Minimum Δ_{BG} depends on θ_{E_6}
 - ▶ \Rightarrow size of tuning depends on low-energy $U(1)$
- ▶ E.g., $U(1)_\psi, d \approx 0.5$ compared to ≈ 0.4 in $E_6\text{SSM}$
- ▶ \Rightarrow lower bound for given $m_{Z'}$ increased
- ▶ $d \neq 0 \Rightarrow m_{Z'}$ still sets minimal sensitivity

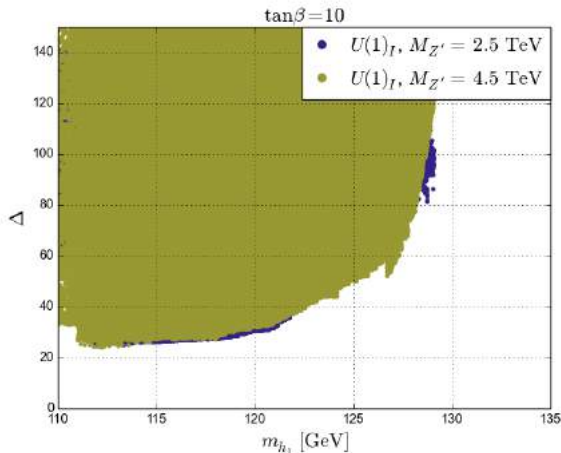
$$U(1)_\eta: Q_\eta = -Q(\theta_{E_6} = \pi - \arctan \sqrt{5/3}), d(\tan \beta = 10) \approx 0.81$$



[arXiv:1503.08929]

- ▶ Largest D -term contribution, $d \approx 0.81$
- ▶ \Rightarrow largest tuning for given $m_{Z'}$
 - ▶ Minimum Δ_{BG} substantially larger than in $U(1)_N, U(1)_\psi$
- ▶ Current $m_{Z'}$ limits \Rightarrow models with large D -terms already "moderately" tuned
- ▶ Severity of tuning depends strongly on θ_{E_6} (i.e. charges)

$$U(1)_I: Q_I = -Q(\theta_{E_6} = \arctan \sqrt{3/5}), d(\tan \beta = 10) \approx -0.01$$



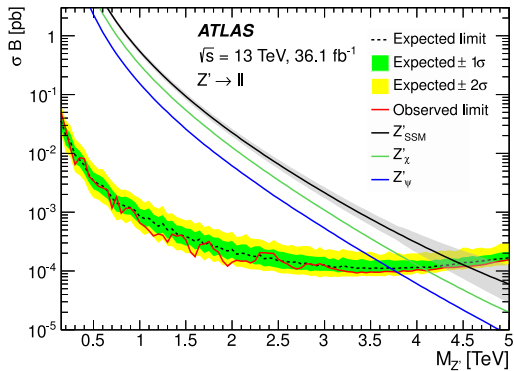
[arXiv:1503.08929]

- ▶ Example of model with suppressed D -terms ($Q_{H_u} = 0$)
- ▶ $\Rightarrow m_{Z'}$ sensitivity much reduced
- ▶ Trade-off: suppresses D -term contribution to m_{h_1} :

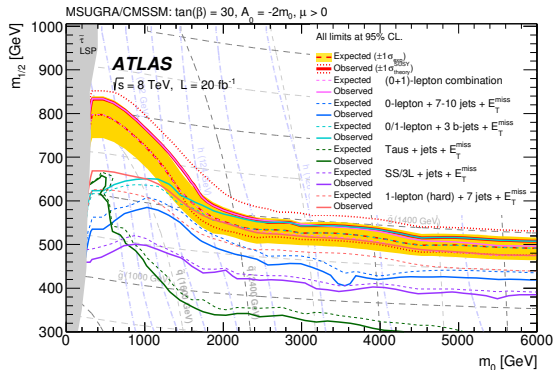
$$m_{h_1}^2 \lesssim m_Z^2 \cos^2 2\beta + \frac{\lambda^2}{2} v^2 \sin^2 2\beta + g_1'^2 v^2 (Q_{H_d} \cos^2 \beta + \underbrace{Q_{H_u} \sin^2 \beta}_{=0})^2$$

- ▶ For current $m_{Z'}$ limits, tuning is MSSM-like

$E_6SSM: M_{Z'} \sim M_S$



[arXiv:1707.02424]



[arXiv:1507.05525]

E_6 SSM: Discrete Symmetries

- ▶ Most general superpotential

$$W \supset g_{ijk}^D \hat{D}_i \hat{Q}_j \cdot \hat{Q}_k + \tilde{g}_{ijk}^E \hat{e}_i^c \hat{D}_j \hat{u}_k^c + y_{ijk}^U \hat{u}_i^c \hat{H}_{uj} \cdot \hat{Q}_k + y_{ijk}^D \hat{d}_i^c \hat{Q}_j \cdot \hat{H}_{dk} + \dots$$

⇒ unacceptable B , L violation and large FCNCs

- ▶ Impose

- ▶ either exact Z_2^L or exact Z_2^B to forbid B , L violating couplings
- ▶ approximate Z_2^H to suppress FCNCs

- ▶ Resulting Higgs, singlet couplings

$$\lambda_{ijk} \hat{S}_i \hat{H}_{dj} \cdot \hat{H}_{uk} \rightarrow \lambda \hat{S} \hat{H}_{d3} \cdot \hat{H}_{u3} + \lambda_{\alpha\beta} \hat{S} \hat{H}_{d\alpha} \cdot \hat{H}_{u\beta} + \tilde{f}_{\alpha\beta} \hat{S}_\alpha \hat{H}_{d\beta} \cdot \hat{H}_{u3} + f_{\alpha\beta} \hat{S}_\alpha \hat{H}_{d3} \cdot \hat{H}_{u\beta}$$

⇒ LSP, NSLP is a (ruled out) “inert” neutralino

- ▶ Simplest viable models impose *another* exact Z_2^S
- ▶ None of these Z_2 symmetries commute with E_6

The SE₆SSM

- ▶ E₆ inspired model arising from 5D or 6D orbifold GUT [2]
- ▶ Complete **27**-plets supplemented by components of **extra 27'-, $\overline{27}'$** -plets
- ▶ Stabilise Higgs potential \Rightarrow pure singlet $\hat{\phi}$
- ▶ $U(1)_{\psi} \times U(1)_{\chi} \rightarrow U(1)_N \times Z_2^M$ at intermediate scale \Rightarrow automatically conserved Z_2^M
- ▶ Dangerous B, L violating operators, FCNCs forbidden by **single exact \tilde{Z}_2^H**

$$\begin{aligned} W_{\text{SE}_6\text{SSM}} = & \lambda \hat{S}(\hat{H}_d \cdot \hat{H}_u) - \sigma \hat{\phi} \hat{S} \hat{S} + \frac{\kappa}{3} \hat{\phi}^3 + \frac{\mu}{2} \hat{\phi}^2 + \Lambda_F \hat{\phi} + \lambda_{\alpha\beta} \hat{S}(\hat{H}_\alpha^d \cdot \hat{H}_\beta^u) \\ & + \kappa_{ij} \hat{S} \hat{D}_i \hat{D}_j + \tilde{f}_{i\alpha} \hat{S}_i(\hat{H}_u \cdot \hat{H}_\alpha^d) + f_{i\alpha} \hat{S}_i(\hat{H}_\alpha^u \cdot \hat{H}_d) + g_{ij}^D(\hat{Q}_i \cdot \hat{L}_4) \hat{D}_j \\ & + h_{i\alpha}^E \hat{e}_i^c(\hat{H}_\alpha^d \cdot \hat{L}_4) + \mu_L(\hat{L}_4 \cdot \hat{L}_4) + \tilde{\sigma} \hat{\phi}(\hat{L}_4 \cdot \hat{L}_4) + W_{\text{MSSM}}(\mu = 0) \end{aligned}$$

EWSB in the SE_6SSM

- ▶ At physical minimum,

$$\langle H_d^0 \rangle = \frac{v_1}{\sqrt{2}}, \quad \langle H_u^0 \rangle = \frac{v_2}{\sqrt{2}}, \quad \langle S \rangle = \frac{s_1}{\sqrt{2}}, \quad \langle \bar{S} \rangle = \frac{s_s}{\sqrt{2}}, \quad \langle \phi \rangle = \frac{\varphi}{\sqrt{2}}$$

- ▶ S, \bar{S} develop along nearly D -flat direction $\langle S \rangle = \langle \bar{S} \rangle$ with

$$\langle S \rangle \approx \langle \bar{S} \rangle \sim \frac{M_S}{\sigma}$$

- ▶ Small $\sigma \Rightarrow M_{Z'}^2 \sim g_1'^2 Q_S^2 (s_1^2 + s_2^2)$ far heavier than M_S
- ▶ \therefore can have $M_{Z'}$ far above limits while keeping sparticles (relatively) light
- ▶ D -term contribution to EW scale also suppressed:

$$c \frac{M_{Z'}^2}{2} \approx -\frac{\lambda^2 s_1^2}{2} + \frac{m_{H_d}^2 - m_{H_u}^2 \tan^2 \beta}{\tan^2 \beta - 1} + d \frac{g_1'^2 Q_S^2 (s_1^2 - s_2^2)}{2}$$

Dark Matter Candidates?

- ▶ Conserved $Z_2^M, \tilde{Z}_2^H \Rightarrow$ two (distinct) DM candidates
- ▶ “Exotics” $\equiv Z_2^E$ odd states, where $\tilde{Z}_2^H = Z_2^M \times Z_2^E$
- ▶ Z_2^E conserved \Rightarrow lightest exotic is stable
- ▶ Limits from exotic Higgs decays, DM direct detection \Rightarrow inert singlinos \tilde{S}_i form subdominant hot DM, $m_{\tilde{S}_i} \ll 1$ eV
- ▶ $M_{Z'} \gg M_S \gg M_Z \Rightarrow$ singlet dominated $\tilde{\chi}^0$'s decouple
- ▶ \Rightarrow account for Ωh^2 with MSSM-like $\tilde{\chi}_1^0$?

	\tilde{Z}_2^H	Z_2^M	Z_2^E
$\hat{Q}_i, \hat{u}_i^c, \hat{d}_i^c, \hat{L}_i, \hat{e}_i^c, \hat{N}_i^c$	-	-	+
$\hat{H}_\alpha^u, \hat{H}_\alpha^d, \hat{S}_i, \hat{D}_i, \hat{\bar{D}}_i$	-	+	-
\hat{H}_u, \hat{H}_d	+	+	+
$\hat{S}, \hat{\bar{S}}$	+	+	+
$\hat{L}_4, \hat{\bar{L}}_4$	+	-	-

Enlarged (8×8) $\tilde{\chi}^0$ sector:

$$M_{\tilde{\chi}^0} = \begin{pmatrix} A & C^T \\ C & B \end{pmatrix}$$

The CSE₆SSM

- ▶ **General model is complicated**
 - ▶ $O(200)$ new parameters (assuming no new sources of CP-violation)
 - ▶ Many masses and mixings
- ▶ Consider constrained model (CSE₆SSM) inspired by gravity mediated SUSY breaking
- ▶ Universal soft breaking parameters: $M_{1/2}, A_0, B_0, m_0$
- ▶ Interested in mechanism decoupling Z' from EWSB conditions \Rightarrow can have large $s = \sqrt{s_1^2 + s_2^2}$
- ▶ Higgsino mass set by $\mu_{\text{eff}} = \lambda s_1 / \sqrt{2} \Rightarrow$ acceptable LSP mass (\lesssim TeV) for small λ
- ▶ \Rightarrow other exotic couplings must be small, otherwise exotic states are tachyonic

Parameter Space Scans

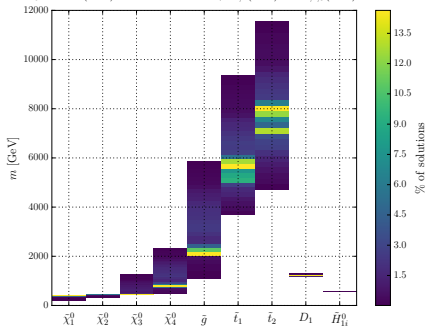
- ▶ Focus on heavy Z' , $s = 650$ TeV, choose fixed μ_{eff}
- ▶ Achieve using semi-analytic solutions for soft parameters:

$$M_i(Q) = p_i(Q)M_{1/2} + q_i(Q)A_0, \quad A_i(Q) = e_i(Q)A_0 + f_i(Q)M_{1/2},$$
$$m_i^2(Q) = a_i(Q)m_0^2 + b_i(Q)M_{1/2}^2 + c_i(Q)A_0M_{1/2} + d_i(Q)A_0^2, \dots$$

- ▶ Fix m_0 from EWSB, $m_0^2 \sim -\frac{b_{H_u}}{a_{H_u}}M_{1/2}^2 - \dots$
- ▶ Implemented in FlexibleSUSY for **full 1-loop masses and 2-loop RGEs**
 - ▶ Resulting “semi-analytic solver” forms part of FlexibleSUSY 2.0 [3], along with many other updates.
- ▶ Require $\Omega h^2 \leq 0.1187$ (micrOMEGAs) and $m_{h_1} = 125.09 \pm 3$ GeV
- ▶ Compare with CMSSM for $|\mu| \sim 400$ GeV and $|\mu| \sim 1$ TeV

Sparticle Mass Spectrum

CSE₆SSM: $\lambda(M_X) = 9.15181 \times 10^{-4}$, $\lambda_{1,2}(M_X) = \kappa_{1,2,3}(M_X) = 10^{-3}$



▶ EWSB conditions $\Rightarrow m_0 > M_{1/2}, A_0$

▶ \therefore MSSM sfermions out of reach of run II

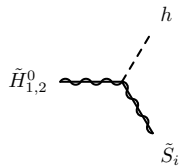
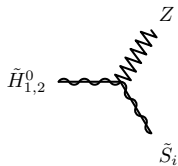
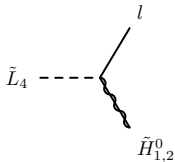
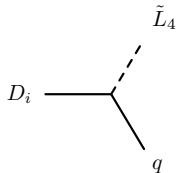
▶ Light exotic fermions can be observable

▶ Exotic leptoquarks D_i :

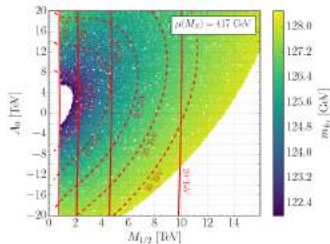
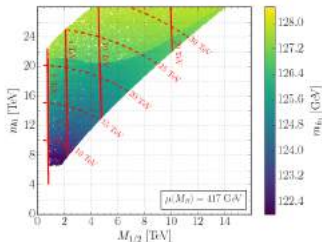
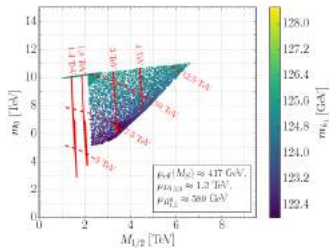
$$pp \rightarrow t \bar{t} \tau^+ \tau^- + E_T^{miss} + X, pp \rightarrow b \bar{b} + E_T^{miss} + X$$

▶ Charged, neutral inert Higgsinos:

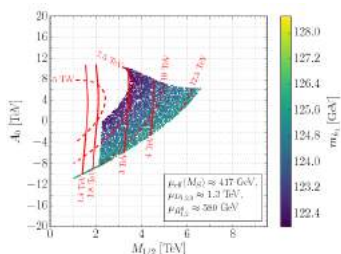
$$pp \rightarrow W W / Z Z / W Z + E_T^{miss} + X$$



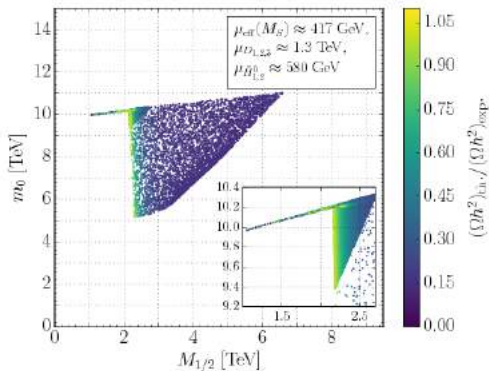
$\mu_{\text{(eff.)}} \approx 400$ GeV: Parameter Space Restrictions



- ▶ Successful EWSB + $m_{h_1} \approx 125$ GeV \Rightarrow large $m_0 > M_{1/2}, A_0$
- ▶ $m_{h_1} \approx 125$ GeV important constraint on range of variation of $M_{1/2}, A_0$
 - ▶ Additional constraints in CSE₆SSM from tachyonic CP-even and CP-odd Higgs states
- ▶ Sfermions heavy, but gluino and EW-inos can be observable

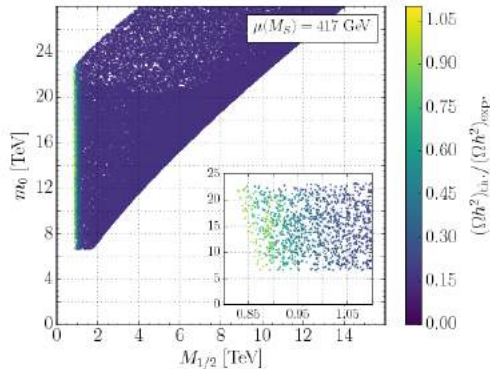


$\mu_{\text{eff.}} \approx 400$ GeV: Relic Density



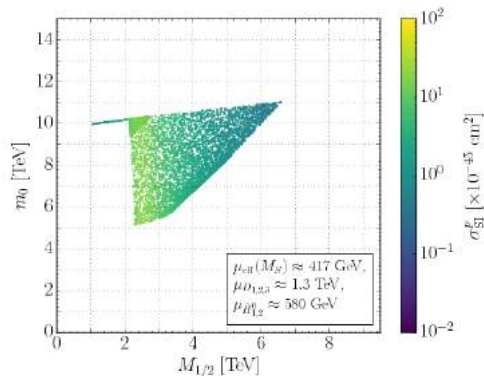
- ▶ $\Omega h^2 \approx 0.1187 \Rightarrow$ “well-tempered” bino-Higgsino $\tilde{\chi}^0$ ($\mu_{\text{eff.}} \sim M_1$)
- ▶ Pair annihilations $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \bar{f} f$

farXiv:1610.033741

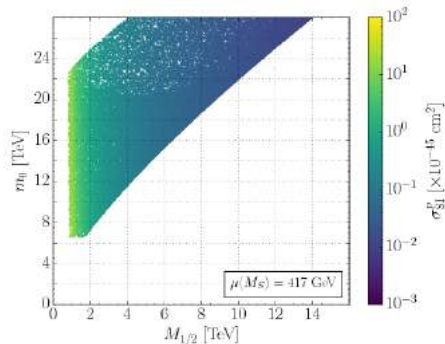


- ▶ MSSM-like nature of neutralino sector clear – almost identical behaviour
- ▶ Existence of A-funnel at $\tan \beta = 10$ notable difference in CSE₆SSM

$\mu_{\text{eff.}} \approx 400$ GeV: Direct Detection Cross Section



- ▶ Well-tempered $\tilde{\chi}_1^0 \Rightarrow$ large mixing, σ_{SI} exceeds, e.g., 90% LUX limits
- ▶ $M_1 \gg \mu_{\text{eff.}}$ both mixing and number density suppressed

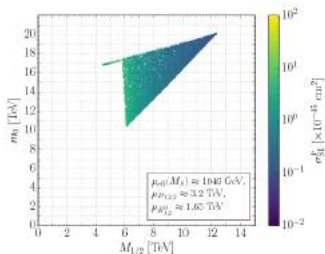
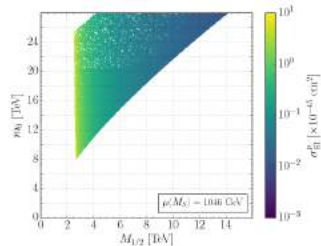
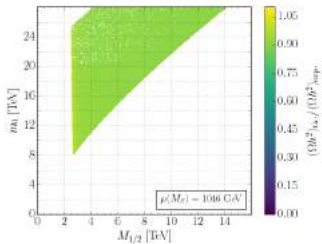
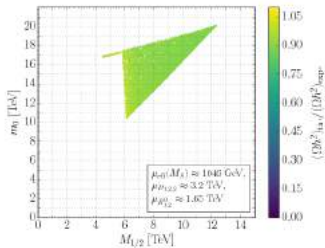


- ▶ σ_{SI} set by $g_{h_1\chi_1\chi_1}$ (t -channel h exchange):

$$g_{h_1\chi_1\chi_1} \approx \frac{1}{2} \left(\sqrt{\frac{3}{5}} g_1 N_{14} - g_2 N_{13} \right) [N_{11}(U_h)_{11} - N_{12}(U_h)_{12}]$$

[arXiv:1610.03374]

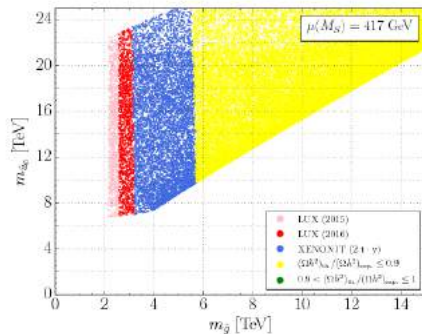
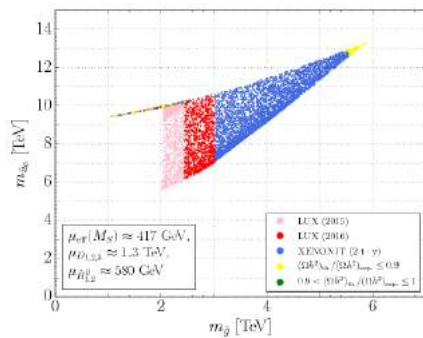
$\mu_{\text{(eff.)}} \approx 1 \text{ TeV}$: Pure Higgsino DM Candidate



- ▶ Similar parameter space constraints due to m_{h_1} , tachyonic states
- ▶ Suppress \tilde{B} fraction \Rightarrow large $M_{1/2}$, $m_{\tilde{g}} \gtrsim 4 \text{ TeV}$, $m_{\tilde{q}} \gtrsim 10 \text{ TeV}$
- ▶ Exotics not forced to be heavy
- ▶ σ_{SI} (mostly) acceptably small

Current and Future Limits

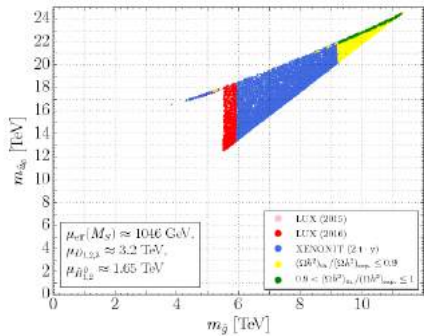
[arXiv:1610.03374]



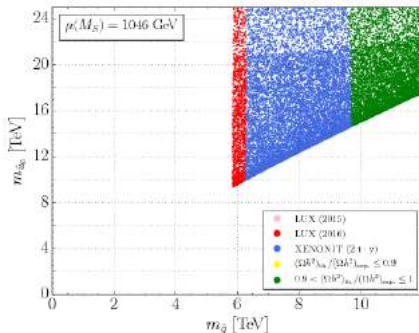
- ▶ LUX \Rightarrow already stringent limits on highly mixed scenarios
- ▶ XENON1T expected to cover many of the remaining solutions

- ▶ SD limits (LUX, IceCube) can also be relevant
- ▶ A-funnel solutions can survive, but in reach of LHC run II \Rightarrow complementarity of searches

Current and Future Limits



- ▶ Non-mixed scenarios expected to be discoverable at XENON1T
- ▶ Larger $M_{1/2}$ and heavy exotics \Rightarrow only accessible at, e.g., LZ



- ▶ Essentially no collider limits for $\mu_{\text{eff.}} \approx 1 \text{ TeV}$ (except possibly exotics)
- ▶ LUX now excludes mixed $\tilde{\chi}_1^0$ even when $m_{\tilde{\chi}_1^0} \approx 1 \text{ TeV}$

Summary

- ▶ E_6 inspired models can address, e.g., little hierarchy problem and μ -problem of MSSM
- ▶ Additional D -terms \Rightarrow also potential source of fine-tuning
- ▶ Bayesian analysis of BSM models automatically incorporates intuitions about naturalness
- ▶ Traditional measures, e.g., Δ_{BG} , agree qualitatively with Bayesian methods
- ▶ Limits on $m_{Z'}$ \Rightarrow conservative limits on Δ_{BG}
 - ▶ Not dependent on assumptions about high-energy model
- ▶ Depends strongly on $U(1)'$, e.g., $U(1)_N$ compared to $U(1)_I$

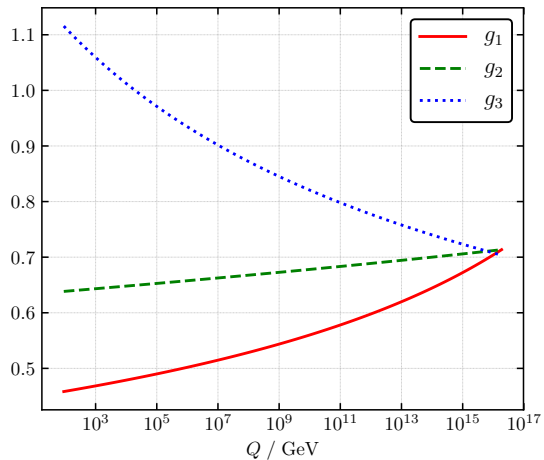
Summary

- ▶ Simplest phenomenologically viable variants, e.g. E_6 SSM, face other problems: multiple discrete symmetries, $m_{Z'} \sim M_S \dots$
- ▶ SE_6 SSM is a well-motivated extension with an exact custodial symmetry
- ▶ DM relic density can be fitted by MSSM-like $\tilde{\chi}_1^0$
- ▶ Direct detection searches \Rightarrow stringent limits on allowed mixing
 - ▶ Mixed states fully accounting for Ωh^2 excluded
- ▶ XENON1T expected to probe much of remainder space, complementary probe to LHC searches
- ▶ E_6 exotics can be light \Rightarrow important possible discovery channel

Thank you for listening!

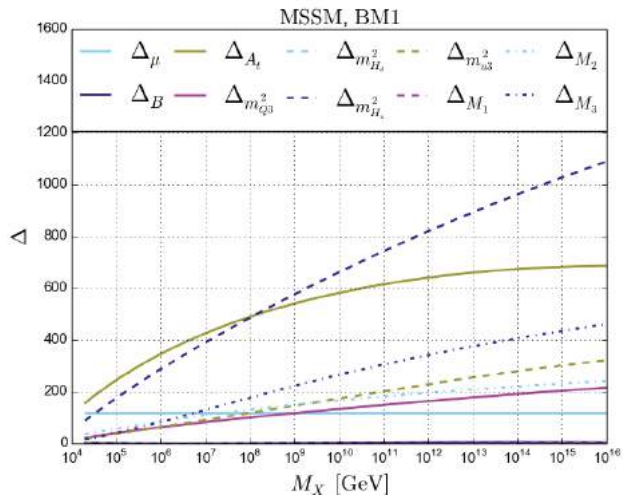
Additional Slides

CMSSM Gauge Coupling Running



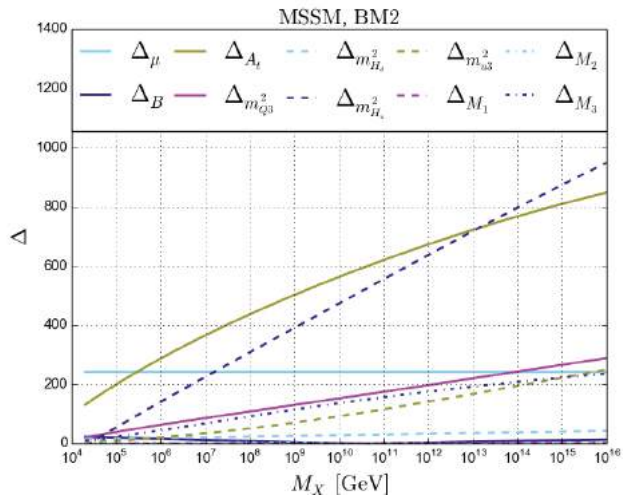
[<http://flexiblesusy.hepforge.org/images.html>]

MSSM BM1: Tuning Breakdown



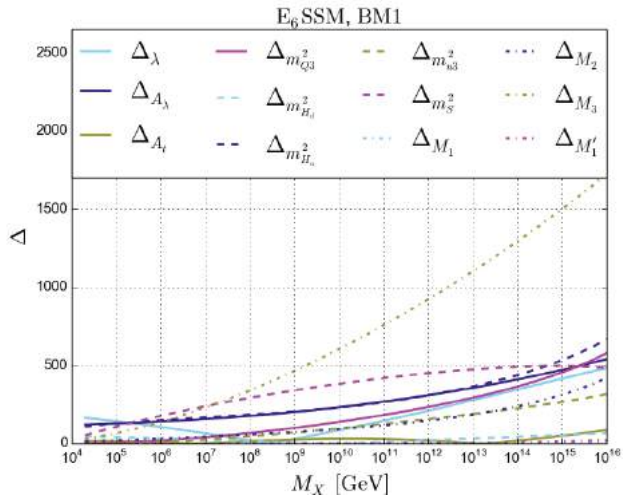
[arXiv:1503.08929]

MSSM BM2: Tuning Breakdown



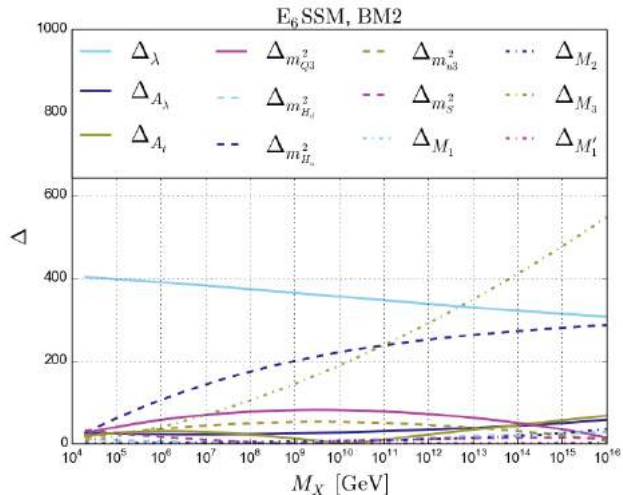
[arXiv:1503.08929]

E₆SSM BM1: Tuning Breakdown



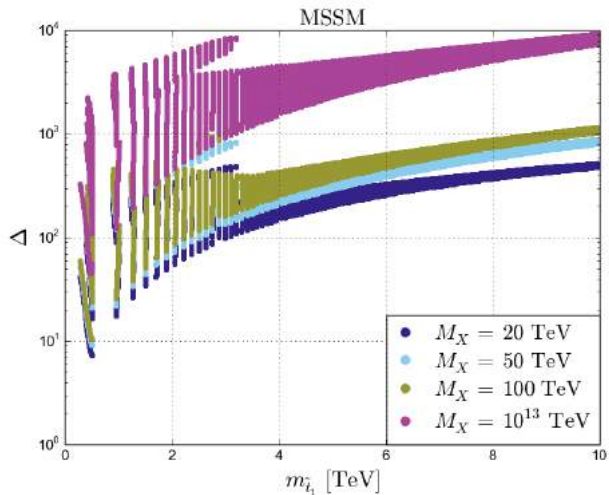
[arXiv:1503.08929]

E₆SSM BM2: Tuning Breakdown



[arXiv:1503.08929]

Low $M_X \Rightarrow$ Low \tilde{t} Tuning



[arXiv:1503.08929]

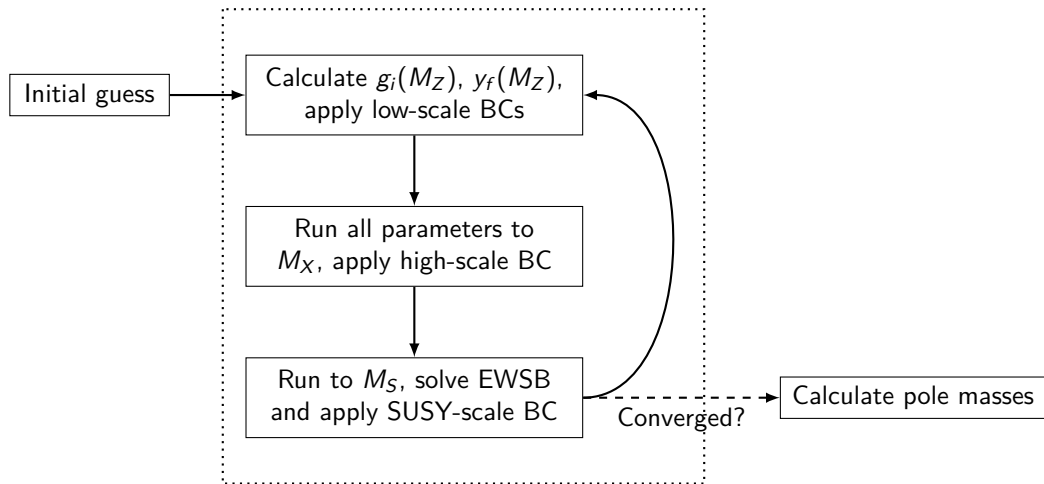
Semi-analytic Solutions

- ▶ Useful technique for studying constrained models (e.g. CE₆SSM)
- ▶ RGEs for SUSY parameters independent of soft SUSY breaking parameters (up to effects of threshold corrections)
- ▶ RGEs for SUSY breaking parameters are system of linear ODEs with variable coefficients
- ▶ Combine with boundary conditions \Rightarrow semi-analytic solutions,

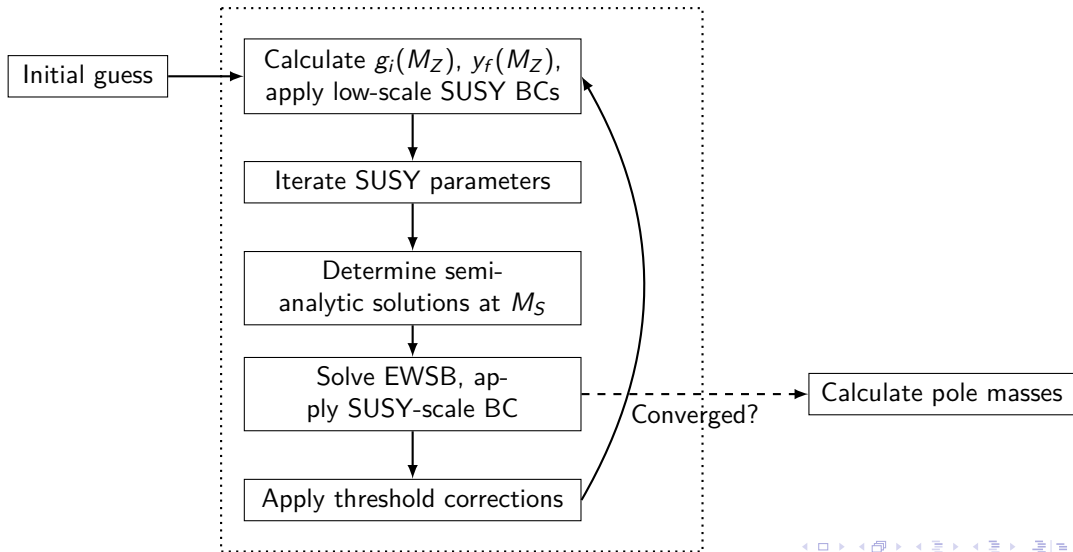
$$M_i(Q) = p_i(Q)M_{1/2} + q_i(Q)A_0, \quad A_i(Q) = e_i(Q)A_0 + f_i(Q)M_{1/2},$$
$$m_i^2(Q) = a_i(Q)m_0^2 + b_i(Q)M_{1/2}^2 + c_i(Q)A_0M_{1/2} + d_i(Q)A_0^2, \dots$$

- ▶ Coefficients depend only on SUSY parameters, calculated numerically
- ▶ Relate SUSY scale parameters to high-scale inputs, e.g. trade m_0 for μ_{eff}
 - ▶ Contrast with traditional approach in CMSSM: μ output for given m_0
 - ▶ N.B. all results here are for fixed μ

Two-scale Algorithm

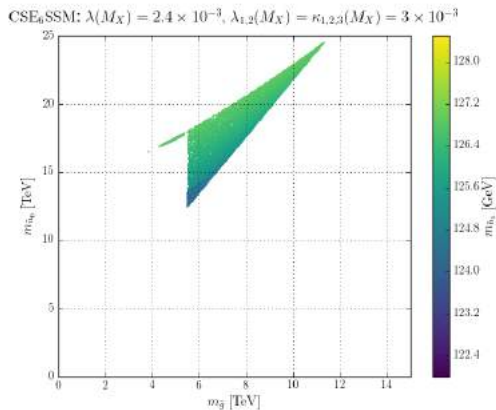


Semi-analytic Algorithm



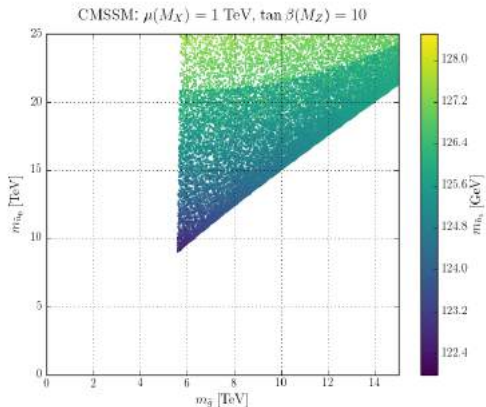
CP-even Higgs Mass in the CSE₆SSM

- ▶ Split spectrum \Rightarrow large logarithms contribute
- ▶ Exotic contributions small, mainly \tilde{t} 's
- ▶ Use EFT calculation of m_{h_1} (SUSYHD)
 - ▶ cross-checked with prototype FlexibleEFTHiggs



[arXiv:1610.03374]

$m_{h_1} \lesssim 130$ GeV
(i.e. bounded from above, as usual)



A-funnel in the CSE₆SSM

- ▶ Solutions at lower $M_{1/2}$ due to $m_{A_1} \sim 2m_{\tilde{\chi}_1^0}$
- ▶ CMSSM: A-funnel requires $\tan \beta \gtrsim 40$
- ▶ CSE₆SSM: tune A_0 for given $\tan \beta$, $M_{1/2}$ to that $m_{A_1} \rightarrow 0$, keeping $m_{\tilde{\chi}_1^0} \sim$ fixed
- ▶ Lightest state A_1 mixture of singlets for $s \gg M_S \gg v$ ($\tan \delta \approx \frac{s_1 s_2}{\varphi \sqrt{s_1^2 + s_2^2}}$):

$$m_{A_1}^2 \approx \cos^2 \delta \left(-2B\mu - 3\frac{\kappa A_\kappa}{\sqrt{2}}\varphi - \sqrt{2}\xi\frac{\Lambda}{\varphi} + \frac{9}{2}\sigma\kappa s_1 s_2 + 2\sqrt{2}\frac{\sigma\mu s_1 s_2}{\varphi} + \frac{\sigma s_1 s_2 \Lambda}{\varphi^2} \right)$$

