LARGE FLUCTUATIONS AND EXTREME EVENTS, DRESDEN, OCTOBER 2015

Power Laws and Criticality



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Power Laws and Criticality, Dresden 2015

1. The size of earthquakes (and other natural hazards)

2. Properties of power-law distributions

• Scale invariance. Divergence of moments

3. Models for criticality

- Galton-Watson model
- Extinction probability
- Size distribution

4. Self-organization towards criticality

- Self-organized branching process
- Manna model, Bak-Tang-Wiesenfeld sandpile model
- Spring-block earthquake models

5. Fitting and goodness-of-fit testing of power-law distributions

Only fools and charlatans predict earthquakes. C. F. RICHTER

Gutenberg-Richter Law (1941)

 Most important law of statistical seismology and a paradigm of complex-systems geophysics

For each earthquake with magnitude $m \geq 4$ there are about

- \star 0.1 with $m\!\geq\!5$
- \star 0.01 with $m\!\geq\!6$, etc...

Number of earthquakes with magnitude $\geq m$

$$N(m) \propto 10^{-bm}$$
, with $b \simeq 1$

Good news! \mapsto Many small earthquakes, few big ones



seismo.berkeley.edu

Gutenberg & Richter, BSSA 1944



• Example: worldwide earthquakes (one-year average)

Kanamori & Brodsky, Rep Prog Phys 2004

 $N(m) \propto 10^{-bm} \Rightarrow \log N(m) = \text{constant} - bm$

Exponential Distribution of Earthquake Magnitudes

• Complementary cumulative distribution (survivor) function

$$S_m(m) = P\{\text{magnitude} \ge m\} | \Rightarrow S_m(m) \propto N(m) \propto 10^{-bm} |$$

• Probability density

$$D_m(m) = \frac{P\{m \le \text{magnitude} < m + dm\}}{dm} = -\frac{dS_m(m)}{dm}$$

verifies $\int_0^\infty D_m(m) dm = 1$ and usually has units! \Rightarrow It is not a probability

Gutenberg-Richter law
$$\Rightarrow D_m(m) \propto 10^{-bm}$$

 $S_m(m)$ and $D_m(m)$ are "the same" only for the exponential distribution A statistician would stop here, wouldn't she?

Which is the Meaning of the Gutenberg-Richter Law?

It depends (of course) on the meaning of magnitude...
 But magnitude is not a proper physical variable (it has no units)!

Moreover: magnitudes reflect radiation only from subportions of the rupture, and they saturate above certain size, rather than giving a physical characterization of the entire earthquake source Ben-Zion, Rev Geophys 2008

• Radiated energy is supposed to be an exponential function of magnitude

$E \propto 10^{3m/2}$

(with proportionality factor around 60 kJ)

An increase of 1 unit in m leads to a factor $\sqrt{10^3}\,{\simeq}\,32$ in E

 \Rightarrow An earthquake with m = 9 is "equivalent" to 1000 of m = 7

• Then, the Gutenberg-Richter law, in terms of $E \propto 10^{3m/2}$:

$$S_m(m) \propto 10^{-bm} \quad \Rightarrow \quad S_E(E) \propto \frac{1}{E^{2b/3}} \simeq \frac{1}{E^{0.7}}$$

Do you know how to perform the change of variables for the density?

$$D_m(m) \propto 10^{-bm} \quad \Rightarrow \quad D_E(E) \propto \frac{1}{E^{\beta}}$$

with

$$\beta = 1 + \frac{2b}{3} \simeq 1.7$$

⇒ Earthquake energy is power-law distributed

Wadati 1932, see Utsu, PAGEOPH 1999

 \Rightarrow Power-law fit cannot be rejected

Main et al. Nature Geosci. 2008





• Compare world with Southern California



Valid up to $m \simeq -4$ in very small regions

Kwiatek et al., Bull Seis Soc Am 2010

• Even valid for fractures in the lab

Baró et al., Phys Rev Lett 2013



In nanofractures valid up to $m\simeq -13$

Åström et al. Phys Lett A 2006

⇒ Enormous range of validity of the Gutenberg-Richter law

• This law is amazing! How can the dynamics of all the elements of a system as complicated as the crust of the earth, with mountains, valleys, lakes, and geological structures of enormous diversity, conspire, as by magic, to produce a law with such extreme simplicity?

P. Bak, 1996

• Other examples of power-law distributions in natural hazards



Forest fires, Malamud et al. Science 1998; PNAS 2005



• Volcanic eruptions, Lahaie & Grasso, J Geophys Res 1998

Auroras, Uritsky et al. J Geophys Res 2002

Freeman & Watkins, Science 2002



• Tsunamis,

Burroughs & Tebbens, PAGEOPH 2005



• Rainfall: flow of water in one point along duration of rain



Peters et al. Phys Rev Lett 2002; J Stat Mech 2010

• Biological extinctions:

Extinction measured as the percentage of extinct families in fixed periods of time (4 millions years)

Sepkoski, Raup; after Bak 1996





Scaling laws never happen by accident G. I. BARENBLATT, 2003

Is there anything special about power-law distributions?

Scale transformation

• Consider a function D(x). Let us perform a linear transformation of the axes

$$\top [D(x)] = c_y D(x/c_x)$$

with $c_i > 0$, for i = x, y. If $c_i > 1$ then \top acts as a mathematical microscope

• For example:

looking at D(x) at the scale of m, $c_x = c_y = 100$ show D(x) at the scale of cm

• Visual example: $\top [D(x)] = c_y D(x/c_x)$ with $c_x = 10$ and $c_y = 2$



• Visual example: $\top [D(x)] = c_y D(x/c_x)$ with $c_x = 10$ and $c_y = 2$



Scale invariance

 Mathematicians are allowed to ask themselves "silly" questions: Invariance under a scale transformation?

$$\top [D(x)] = c_y D(x/c_x) = D(x)$$

Solution?

• The only solution of $D(x) = c_y D(x/c_x)$ for all c_x is the power law

$$D(x) \propto rac{1}{x^{eta}}$$
 with $eta = -rac{\ln c_y}{\ln c_x}$ i.e., $c_y = rac{1}{c_x^{eta}}$

Direct substitution confirms that it is a solution indeed

• Example: $D(x) \propto \sqrt{x}$ (i.e., $\beta = -1/2$). If $c_x = 10 \Rightarrow c_y = \sqrt{10}$



Difference between $\beta < 0$ (increasing power law) and $\beta > 0$ (decreasing)

★ if $\beta < 0$ and $c_x > 1$ then $c_y = 1/c_x^\beta > 1$ ★ if $\beta > 0$ and $c_x > 1$ then $c_y = 1/c_x^\beta < 1$

Demonstration •

Differentiate both sides of $D(x) = c_y D(x/c_x)$ with respect x and isolate c_y

$$\frac{D'(x)}{D'(x/c_x)/c_x} = c_y = \frac{D(x)}{D(x/c_x)}$$

so, separating variables x and x/c_x and multiplying by x

$$\frac{xD'(x/c_x)}{c_xD(x/c_x)} = \frac{xD'(x)}{D(x)}$$

which has to be valid for all $c_{\mathcal{X}}$, so, it only can be a constant (+, -, or 0),

$$\frac{xD'(x)}{D(x)} = \text{ constant} = -\beta \qquad \Rightarrow \qquad D(x) \propto \frac{1}{x^{\beta}} \qquad \text{for } x > 0$$

Meaning of scale invariance?

• Power-law distributions do not have a characteristic scale

One can define the time unit (or a clock) from the law of radioactive decay (which is an exponential, not a power law)

But one cannot define a unit of distance from the law of gravitation (which is a power law)

In the same way that one cannot built a compass from a sphere (which has rotational symmetry)

- So, earthquake energies have no characteristic scale
- \Rightarrow It is not possible to answer this simple question:

"How big are earthquakes in this region?"

Implications for extreme events

• We have already seen that the GR law for earthquakes implies that:

large earthquakes do not play a special role, they follow the same law as small earthquakes

 \Rightarrow general theory encompassing all earthquakes, large and small P. Bak, 1996

• But scale invariance goes beyond this fact:

there is no unarbitrary way to separate ordinary events from extreme events

(at least attending the statistics of event sizes)

Discrete scale invariance

• We can consider the constant eta as a complex number, $eta
ightarrow eta - \omega i$

$$\Rightarrow \frac{1}{x^{\beta}} \to x^{-\beta + \omega i} = x^{-\beta} e^{i\omega \ln x} \qquad \text{and substitute in } c_y D(x/c_x) = D(x)$$

Then, if c_x is real, then $c_y = 1/c_x^{\beta-i\omega} = c_x^{-\beta} e^{i\omega \ln c_x}$ is complex (in general) Imposing that c_y is positive real $c_x = \exp(2\pi n/\omega)$ with $n = 0, \pm 1, \pm 2 \dots$

Thus, scale invariance does not hold for all c_x but for discrete values In this case, the real part and the imaginary part are also scale invariant

$$\operatorname{Re}[x^{-\beta+\omega i}] = \frac{1}{x^{\beta}}\cos(\omega \ln x)$$
 or $\operatorname{Im}[x^{-\beta+\omega i}] = \frac{1}{x^{\beta}}\sin(\omega \ln x)$

Scale invariance for multivariate functions

• Consider D(x, y) and a scale transformation $\top [D(x, y)] = c_z D(x/c_x, y/c_y)$ The scale-invariance condition $D(x, y) = c_z D(x/c_x, y/c_y)$ has a unique solution

$$D(x,y) = x^{-\beta} F(y/x^{\alpha})$$
 for all $c_x > 0$

which is called a scaling law, with

$$c_{\mathcal{Y}} = c_{x}^{lpha}$$
 and $c_{z} = rac{1}{c_{x}^{eta}}$

and the scaling function? F() is arbitrary

- Equivalent expressions: $D(x,y) = x^{-eta}F_2(x/y^{1/lpha}) = y^{-eta/lpha}F_3(x/y^{1/lpha})$, etc.
- We will distinguish scaling laws from power laws For univariate functions both are the same, with F = constant

Christensen & Moloney 2005

Mean earthquake energy...?

$$E[E] = \langle E \rangle = \int_{min}^{\infty} ED(E) dE \propto \int_{min}^{\infty} \frac{dE}{E^{\beta - 1}} = \infty$$

- ... is infinite! (because $1 < \beta \leq 2$)
- Higher-order moment are also infinite.
- Which is the problem? Is mathematical?

This process has a mean waiting time between events which is infinite:

 $t_{i+1} = t_i + (1 - u_i)^{1/(\beta - 1)}$ with u_i uniform random in [0, 1)

Is physical then? The Earth contains a finite amount of energy!

• What does $\langle E \rangle = \infty$ mean in practice?

• Consider the average up to the N-th event, $\bar{E} = (E_1 + E_2 + \cdots + E_N)/N$



The rare big events are crucial for energy dissipation \mapsto Bad news!!!

Discrete analog: the St. Petersburg paradox

N. Bernoulli 1713 & D. Bernoulli 1738

• Consider a game of chance in which a player tosses a (fair) coin until a tail appears for the 1st time. Each toss doubles the payoff

Outcome	Probability	Payoff
tail	$p_1 = 1/2$	1\$
heads,tail	$p_2 = 1/4$	2 \$
heads,heads, <mark>tail</mark>	$p_3 = 1/8$	4 \$
	1	7 1
heads heads, <mark>tail</mark>	$p_k = 1/2^{\kappa}$	2^{k-1} \$ for k tosses, in general

• You are a casino: which would be the fair price to pay to enter the game?

$$\langle \mathsf{payoff} \rangle = \sum_{k=1}^{\infty} p_k \times \mathsf{payoff}(k) = \sum_{k=1}^{\infty} \frac{1}{2^k} \times 2^{k-1} \$ = \frac{1}{2} \times 1\$ + \frac{1}{4} \times 2\$ + \ldots = \infty$$

• Note that the duration k of the game is geometrically (exponentially) distributed

$$p_k = \frac{1}{2^k} = e^{-k \ln 2} = 10^{-k \log 2} \implies \langle K \rangle = \sum_{k=1}^{\infty} p_k k = \frac{1}{1/2} = 2$$

so, the duration of the game is analogous to magnitude, with $b = \log 2 \neq 1$

- But the payoff $= 2^{k-1} \propto 10^k \log 2 = 10^{ck}$ is analogous to energy, with $c = \log 2$
- Then, the payoff follows a (sort of) discrete power-law distribution with

$$\beta = 1 + \frac{b}{c} = 1 + \frac{\log 2}{\log 2} = 2$$

This is in the "boundary" of having a finite mean

Laplace transform

• Consider $D_x(x)$ defined for $x \ge 0$, then

$$ilde{D}_{x}(z)=\int_{0}^{\infty}e^{-zx}D_{x}(x)dx=\langle e^{-zX}
angle$$

if $D_{\mathcal{X}}(x)$ is a probability density, normalization implies $\tilde{D}_{\mathcal{X}}(z=0)=1$

• Assuming that $\tilde{D}_x(z)$ exists and that all moments $\langle X^n \rangle$ are finite, and using $e^{-zx} = \sum_{n=0}^{\infty} (-1)^n z^n x^n / n!$

$$\tilde{D}_x(z) = 1 - \langle X \rangle z + \frac{1}{2} \langle X^2 \rangle z^2 - \ldots = \sum_{n=0}^{\infty} (-1)^n \frac{\langle X^n \rangle z^n}{n!}$$

so, the Laplace transform of $D_x(x)$ is a sort of moment generating function

Sum and rescaling of independent random variables

• Define S = X + Y, then $F_s(s) = P\{ sum < s\} = P\{Y < s - X\} \Rightarrow$

$$F_{s}(s) = \int_{0}^{s} dx \int_{0}^{s-x} dy D_{x}(x) D_{y}(y) = \int_{0}^{s} dx D_{x}(x) F_{y}(s-x)$$

Differentiating with the Leibniz rule, $D_{m{s}}(s)=dF_{m{s}}(s)/ds=$

$$= \int_0^s dx D_x(x) \frac{dF_y(s-x)}{ds} + D_x(x) F_y(s-x) \Big|_{x=s} = \int_0^s dx D_x(x) D_y(s-x)$$

Calculating the Laplace transform, with $\theta(x)$ the step function,

$$\tilde{D}_s(z) = \int_0^\infty ds \, e^{-zs} D_s(s) = \int_0^\infty ds \, e^{-zs} \int_{-\infty}^\infty dx D_x(x)\theta(x) D_y(s-x)\theta(s-x)$$

$$\Rightarrow \tilde{D}_s(z) = \int_0^\infty dy \, e^{-zy} D_y(y) \int_0^\infty dx \, e^{-zx} D_x(x) = \tilde{D}_x(z) \tilde{D}_y(z) \quad \text{using } s - x = y$$

The sum is a convolution of D_x and D_y , which turns a product of $ilde{D}_x$ and $ilde{D}_y$

skip!

Sum and rescaling of independent random variables

• Define S = X + Y, then, the Laplace transform of the distribution of S

$$\tilde{D}_s(z) = \langle e^{-zS} \rangle = \int_0^\infty ds \, e^{-zs} D_s(s) = \int_0^\infty \int_0^\infty dx dy D_x(x) D_y(y) \, e^{-z(x+y)}$$

where we have used independence $[D_{x,y}(x,y) = D_{x}(x)D_{y}(y)]$, then

$$\tilde{D}_{s}(z) = \int_{0}^{\infty} dx D_{x}(x) e^{-zx} \int_{0}^{\infty} dy D_{y}(y) e^{-zy} = \tilde{D}_{x}(x) \tilde{D}_{y}(y)$$

So, the Laplace transform of the sum is the product of $\tilde{D}_x(z)$ and $\tilde{D}_y(z)$ It is not necessary to know that $D_s(s)$ is the convolution of $D_x(x)$ and $D_y(y)$

In general, if $S = X_1 + X_2 + \dots + X_N$, then •

$$ilde{D}_{s}(z)=\left[ilde{D}_{x}(z)
ight]^{N}$$

when all X_i are independent and identically distributed Rescaling by a constant, $R\!=\!S/C$

•

$$\tilde{D}_r(z) = \int_0^\infty dr D_r(r) e^{-zr} = \int_0^\infty ds D_s(s) e^{-z(s/C)} = \tilde{D}_s(z/C)$$

Defining the rescaled mean, or "non-conserved" average •

$$R = \frac{X_1 + X_2 + \dots + X_N}{N^{1/\alpha}} \qquad \Rightarrow \qquad \tilde{D}_r(z) = [\tilde{D}_x(z/N^{1/\alpha})]^N$$

Introducing a cumulant generating function •

$$G_x(z) = \ln \tilde{D}_x(z) \qquad \Rightarrow \qquad G_r(z) = NG_x(z/N^{1/\alpha})$$

Sum of X's turns into product of m.g.f. and into a sum of cumulant g.f. (if independence holds) Note: g.f. = generating function, m.g.f. = moment g.f.

• If the moments are finite (and the generating function exists)

$$\tilde{D}_x(z) = 1 - \langle X \rangle z + \frac{1}{2} \langle X^2 \rangle z^2 - \dots$$

Considering $\ln(1-y) = -y - y^2/2 - y^3/3 - \ldots$, with $-1 \le y < 1$, then, the cumulant generating function

$$G_x(z) = \ln \tilde{D}_x(z) = \ln \left[1 - \left(\langle X \rangle z - \frac{1}{2} \langle X^2 \rangle z^2 + \dots \right) \right] =$$

$$-\left(\langle X\rangle z - \frac{1}{2}\langle X^2\rangle z^2 + \dots\right) - \frac{1}{2}\left(\langle X\rangle z - \dots\right)^2 + \dots = -\langle X\rangle z + \frac{\langle X^2\rangle - \langle X\rangle^2}{2}z^2 - \dots$$

From the coefficients we can obtain the cumulants: $\langle X \rangle$, σ^2 , etc.

Distributions stable under "averaging"

• Again a "silly" question: let us look at the fixed points of this transformation

$$G^*(z) = NG^*(z/N^{1/\alpha})$$

This is the scale invariance condition, whose only solution for all N is

 $G^*(z) \propto z^{\alpha}$

• In the case of the arithmetic mean, lpha=1, then $D_x^*(x)=\delta(x-\mu)$, indeed

$$\tilde{D}_x^*(z) = \int_0^\infty dx \, e^{-zx} \delta(x-\mu) = e^{-\mu z} \qquad \Rightarrow \qquad G_x^*(z) = \ln D_x^*(z) = -\mu z$$

where $\delta(x-\mu)$ is a Dirac delta "function", which has mean μ and zero variance
Domain of attraction of the Dirac delta distribution

• Considering the expansion of $G_x(z)$ into cumulants (if they exist and are finite)

$$G_x(z) = -\langle X \rangle z + \frac{\sigma^2}{2} z^2 - \dots$$

Applying the scale transformation we get the distribution of the mean

$$G_{\bar{x}}(z) = G_{r}(z) = NG_{x}(z/N) = N \left[-\langle X \rangle \frac{z}{N} + \frac{\sigma^{2}}{2} \left(\frac{z}{N} \right)^{2} - \dots \right] \to -\langle X \rangle z$$

The distribution of the mean tends to a delta centered at $\langle X \rangle$ when $N \to \infty$ So, the fixed point is attractive if $G_x(z)$ exists and all moments are finite We will see that the domain of attraction is even bigger

• This constitutes a version of the law of large numbers (weak version)

It is somehow analogous to the central limit theorem also Note that the Gaussian (normal) distribution also tends to a delta (because we do not have zero mean) If we had subtracted the mean the "central limit" would have been Gaussian 38

Stability and domain of attraction for "non-conserved" averaging

• Coming back to the general rescaled mean, $G^*(z) \propto z^{lpha}$, consider lpha = 1/2

$$G_x^*(z) = -2a\sqrt{z} \qquad \Rightarrow \qquad D_x^*(z) = e^{-a^2/x} \frac{a}{\sqrt{\pi x^{3/2}}}$$

Abramowitz & Stegun, 29.3.82; Bouchaud & Georges, *Phys Rep* 1990

As $(x_1 + x_2 + \dots + x_N)/N^2$ converges, the mean diverges linearly with N

- Do it yourself! Simulate N random values of X. How?
 - \star Consider the transformation $X = 1/Y^2$
 - \star Y follows a half-normal (half-Gaussian) distribution
 - \star Use standard algorithm (like Box-Muller transformation) to simulate Y
- Alternative: simulate a power-law with exponent $3/2 \Rightarrow$ What happens?

• Power-law distributions belong to the domain of attraction of $G^*(z) \propto z^{\alpha}$ Consider $D_x(x) = B/x^{1+\rho}$ for $x \ge c > 0$ (and 0 otherwise), then $B = \rho c^{\rho}$ and

$$\tilde{D}_x(z) = B \int_c^\infty e^{-zx} x^{-\rho-1} dx = B z^{\rho} \Gamma(-\rho, cz)$$

with $\Gamma(\gamma,z) = \int_z^\infty u^{\gamma-1} e^{-u} du$ the incomplete gamma function, with expansion

$$\Gamma(\gamma, z) = \Gamma(\gamma) - z^{\gamma} \sum_{n=0}^{\infty} \frac{(-z)^n}{(\gamma+n)n!} \qquad \gamma \neq 0, -1, -2, -3...$$
 Abramowitz & Stegun 6.5.29

with $\Gamma(\gamma) = \Gamma(\gamma, 0)$ for $\gamma > 0$ and $\Gamma(\gamma) = \Gamma(\gamma + 1)/\gamma$ for $\gamma < 0$ (non-integer)

$$\Rightarrow z^{\rho} \Gamma(-\rho, z) = z^{\rho} \Gamma(-\rho) - \sum_{n=0}^{\infty} \frac{(-z)^n}{(n-\rho)n!} = \frac{1}{\rho} \left[\rho \Gamma(-\rho) z^{\rho} + \left(1 + \frac{\rho z}{1-\rho} + \dots \right) \right]$$

 $ho
eq 0, 1, 2, \ldots$ We are interested in $G_x(z) = \ln ilde{D}_x(z) = \ln B z^
ho \Gamma(ho, cz)$, so

$$\ln z^{\rho} \Gamma(-\rho, z) = -\ln \rho + \ln \left[\right] = -\ln \rho + \rho \Gamma(-\rho) z^{\rho} + \frac{\rho z}{1-\rho} + \dots$$

$$\Rightarrow G_x(z) = \ln \frac{B}{c^{\rho}} + \ln c^{\rho} z^{\rho} \Gamma(-\rho, cz) = \rho \Gamma(-\rho) c^{\rho} z^{\rho} + \frac{\rho cz}{1-\rho} + \dots$$

using again the expansion of the logarithm. Applying the transformation

$$G_{T}(z) = NG_{x}(z/N^{1/\alpha}) = \mathbb{N} \left[\rho \Gamma(-\rho) c^{\rho} \left(\frac{z}{N^{1/\alpha}} \right)^{\rho} + \frac{\rho c}{1-\rho} \left(\frac{z}{N^{1/\alpha}} \right) + \dots \right]$$

If $0 < \rho < 1$ and $\alpha = \rho$ then $G_{T}(z) \to \rho \Gamma(-\rho) c^{\rho} z^{\rho}$
If $\rho > 1$ and $\alpha = 1$ then $G_{T}(z) = G_{\overline{x}}(z) \to -\frac{\rho c}{\rho - 1} z$

where the coefficient is the mean of the power-law distribution in this case (in any other case $G_{T}(z)
ightarrow 0$ or ∞)

• The domain of attraction includes distributions that are asymptotically power laws

$$D_X(x)\sim {B\over x^{1+
ho}} \qquad {
m for} \; x
ightarrow\infty$$

with $B \neq \rho c^{
ho}.$ If ho is not a positive integer

$$ilde{D}_x(z) \sim B\Gamma(-
ho)z^
ho + \sum_{n=0}^\infty rac{a_n(-z)^n}{n!}$$

Bleistein & Handelsman 4.6.23

which is, except for the multiplying constants, the same as before (with $a_0 = 1$) So, again, there are 2 cases:

$$\begin{array}{ll} \mbox{If } 0 < \rho < 1 & \mbox{ and } \alpha = \rho & \mbox{ then } & G_{T}(z) \rightarrow B\Gamma(-\rho)z^{\rho} \end{array} \\ \label{eq:gamma-field} \mbox{If } \rho > 1 & \mbox{ and } \alpha = 1 & \mbox{ then } & G_{T}(z) = G_{\bar{x}}(z) \rightarrow a_{1}z \end{array}$$

• Reciprocally, $G_x(z) \propto z^
ho$ corresponds to a distribution that is asymptotically power law if 0 <
ho < 1

Summary

• Assuming independence:

If the moments of $D_x(x)$ are finite and its generating function exist or If $D_x(x)$ is asymptotically a power law with exponent $\beta = 1 + \rho > 2$ \Rightarrow the arithmetic mean \bar{x} follows a Dirac's delta distribution

$$\bar{x} = \frac{1}{N} \sum_{i=1}^{N} x_i \quad \sim \quad \delta(\bar{x} - \langle x \rangle) \quad \text{when } N \to \infty$$

 \Rightarrow Law of large numbers

Feller 1971

 \Rightarrow Makes sense of the arithmetic mean!

• Assuming independence:

If $D_x(x)$ is asymptotically a power law with exponent $\beta = 1 + \rho < 2$ (but >1)

$$\frac{\bar{x}}{N^{1/\rho-1}} = \frac{1}{N^{1/\rho}} \sum_{i=1}^{N} x_i \quad \text{ follows a power-law tailed distribution}$$

with exponent $1+\rho,$ when $N\!\rightarrow\!\infty$

- \Rightarrow The arithmetic mean diverges as $\mathbb{N}^{1/\rho-1}$ (as $1/\rho 1 > 0$)
- \Rightarrow Case of the generalized central limit theorem Bouchaud & Georges, Phys Rep 1990
- Therefore, the sum of earthquake energies "converges", if rescaled by

$$N^{1/\rho} = N^{1/(\beta-1)} = N^{3/(2b)}$$

Standard (conserved) averaging leads to divergence of \bar{E} as $N^{3/(2b)-1} \simeq \sqrt{N}$

• Concrete example, case $\beta = 1 + \rho = 3/2$

$$D_x(x)\!\sim\!rac{A}{x^{3/2}}$$
 for "large" x

The "correct" average to get convergence is

$$\frac{\bar{x}}{N^{1/\rho-1}} = \frac{1}{N^{1/\rho}} \sum_{i=1}^{N} x_i = \frac{1}{N^2} \sum_{i=1}^{N} x_i = y$$

Then, for $N \to \infty,$ the "non-conserved" average y follows

$$D_y(y) = e^{-a^2/y} \frac{a}{\sqrt{\pi} \, y^{3/2}}$$

with $a=\,-\,A\Gamma(-1/2)/2\,{=}\,\sqrt{\pi}A$

What Should One Expect from a Theory of a Complex Phenomenon?

P. Bak, How Nature Works, 1996

• A theory must be abstract

- A theory of life does not need to predict elephants (if your theory predicts elephants it is not general enough)
- ***** Explain why there is variability, or what typical patterns may emerge
- ★ If ... we concentrate on an accurate description of the details, we lose perspective

• A theory must be statistical

- ★ Collecting anecdotal evidence can only be an intermediate goal.
- ★ Anecdotal evidence carries weight only if enough of it can be gathered to form a statistical statement.
- ★ Confrontation between theories and ... observations... takes place by comparing the statistical features of general patterns.

- The abstractness and the statistical, probabilistic nature of any such theory might appear revolting to geophysicists, biologists, and economists, expecting to aim for photographic characterization of real phenomena.
- Perhaps too much emphasis has been put on detailed prediction ... in today's materialistic world.
- To predict the statistics of actual phenomena rather than the specific outcome is a quite legitimate and ordinary way of confronting theory with observations.

I cannot imagine a theory of earthquakes that does not explain the GR law Per Bak, 1996 **Domino-like theory: Otsuka's model** (1971)

Earthquake rupture = cascade process of topplings, but:

(1) No domino effect: one toppling does not lead to another one and so on





M. A. Francisco

(2) Pieces are not in a row, rather, in a network or tree, and disordered

 \Rightarrow When one piece topples, what happens next is random



appadvice.com

Otsuka, Zisin 1971

Seismic fault =

patches that may fail and trigger other patches to fail with some probability and so on Kanamori & Mori, in Boschi et al. 2000

Besides gambling, many probabilists have been interested in reproduction G. GRIMMETT AND D. STIRZAKER, 2001

Galton-Watson (Branching) Process (1873)

• Definition

Start with 1 "element" (parent) which generates K = 0, 1, ... elements (offsprings) with some probability $p_0, p_1, ...$ and so on...

$K\space{-1mu}\sp$



Wikipedia www.wolframalpha.com

Galton was not interested in earthquakes

Rather, he was worried by the extinction of prominent families:

a rise in physical comfort and intellectual capacity is necessarily accompanied by diminution in "fertility"... If that conclusion be true, our population is chiefly maintained though the "proletariat," and thus a large element of degradation is inseparably connected with those elements which tend to ameliorate the race

Extinction

• $N_t =$ total number of elements in generation t (with $N_0 = 1$)



Extinction \Rightarrow $N_t = 0$ at some t

• Extinction = extinction in t = 1 or in t = 2 or $\ldots \in \lim_{t \to \infty} \{N_t = 0\}$

$$\Rightarrow P_{extinction} = \lim_{t \to \infty} P\{N_t = 0\}$$

• Probability generating function of a discrete random variable \boldsymbol{X}

$$f_X(z) = \langle z^X \rangle = \sum_{x=0}^{\infty} P\{X = x\} z^x = P\{X = 0\} + P\{X = 1\} z + \dots$$

 $\Rightarrow f_X(0) = \mathbb{P}\{X = 0\}$

This is valid for any random variable, also for N_t , so,

$$\Rightarrow P_{extinction} = \lim_{t \to \infty} f_{N_t}(0)$$

which is easier to calculate

• Main equation

$$N_{t+1} = \sum_{i=1}^{N_t} K_i(t)$$

• If N_t were a constant

$$f_{N_{t+1}}(z) = [f_K(z)]^{N_t}$$

Proof:

$$f_{N_{t+1}}(z) = \langle z^{N_{t+1}} \rangle = \langle z^{\sum_i K_i} \rangle = \langle z^{K_1} \cdots z^{K_{N_t}} \rangle = \langle z^{K_1} \rangle \cdots \langle z^{K_{N_t}} \rangle = [f_K(z)]^{N_t},$$

assuming independence.

- Let us repeat, $N_{t+1} = \sum_{i=1}^{N_t} K_i(t)$. If N_t is constant, $f_{N_{t+1}}(z) = [f_K(z)]^{N_t}$
- But N_t is random, so

$$f_{N_{t+1}}(z) = f_K^{t+1}(z)$$

with $f_K^{t+1}(z) = f_K(f_K(\dots f_K(z) \dots)) = \text{composition } t+1 \text{ times}$ Proof:

$$f_{N_{t+1}}(z) = \langle z^{N_{t+1}} \rangle = \langle \langle z^{N_{t+1}} \rangle_{K_i} \rangle_{N_t} = \langle [f_K(z)]^{N_t} \rangle_{N_t} = f_{N_t}(f_K(z)).$$

As $f_{N_1}(z) = f_K(z)$, then

$$\Rightarrow f_{N_2}(z) = f_{N_1}(f_K(z)) = f_K(f_K(z)) \equiv f_K^2(z)$$

and the result follows by induction

• In conclusion

$$P_{extinction} = \lim_{t \to \infty} P\{N_t = 0\} = \lim_{t \to \infty} f_{N_t}(0) = \lim_{t \to \infty} f_K^t(0)$$

• Let us calculate $f_{N_t}(z)$. Note that $N_1 = K \Rightarrow f_{N_1}(z) = f_K(z)$. Also

$$N_2 = \sum_{i=1}^{N_1} K_i$$
, and, in general $N_{t+1} = \sum_{i=1}^{N_t} K_i$

If $M = \sum_{i=1}^N K_i$, with N constant, then

$$f_M(z) = \langle z^M \rangle = \langle z^{K_1} \rangle = \langle z^{K_1} \cdots z^{K_N} \rangle = \langle z^{K_1} \rangle \cdots \langle z^{K_N} \rangle = [f_K(z)]^N,$$

assuming independence between the K_i 's. But if N is random, with $f_N(z)=\langle z^N\rangle$, then

$$\begin{split} f_M(z) &= \langle z^M \rangle = \left\langle \langle z^M \rangle_{K_i} \right\rangle_N = \left\langle [f_K(z)]^N \rangle_N = f_N(f_K(z)) \right. \\ \\ \Rightarrow f_{N_2}(z) &= f_{N_1}(f_K(z)) = f_K(f_K(z)) \equiv f_K^2(z) \end{split}$$

In the same way

$$N_{t+1} = \sum_{i=1}^{N_t} K_i$$

As $f_M(z)=f_N(f_K(z)),$ then, $f_{N_3}(z)=f_{N_2}(f_K(z))=f_K^2(f_K(z))\equiv f_K^3(z)$ In general, by induction

$$f_{N_t}(z) = f_{N_{t-1}}(f_K(z)) \equiv f_K^t(z)$$
 $(t - \text{times composition})$

Therefore

$$P_{extinction} = \lim_{t \to \infty} P\{N_t = 0\} = \lim_{t \to \infty} f_{N_t}(0) = \lim_{t \to \infty} f_K^t(0)$$

Expected size of population at t

- Property of $f_X(z) = \sum_{x=0}^{\infty} p_x z^x \Rightarrow f'_X(1) = \langle X \rangle$ Valid for any generating function, so, in the same way $f'_X(1) = \langle X \rangle$ • $f'_{N_t}(1) = \langle N_t
 angle$
- •

$$\frac{df_{N_t}(z)}{dz}\bigg|_{z=1} = \left.\frac{df_K^t(z)}{dz}\bigg|_{z=1} = \left.\frac{df_K(f_K^{t-1}(z))}{dz}\right|_{z=1} = \left.f_K'(f_K^{t-1}(z))\frac{df_K^{t-1}(z)}{dz}\right|_{z=1}$$

by the chain rule, and by induction

$$\langle N_t \rangle = f'_K(f_K^{t-1}(z))f'_K(f_K^{t-2}(z))\cdots f'_K(f_K^2(z))f'_K(f_K(z))f'_K(z)\Big|_{z=1}$$

using $f_K(1) = \mathbf{1} \Rightarrow f_K^2(1) = 1$, etc., and $f'_K(1) = \langle K \rangle$ then $\langle N_t \rangle = \langle K \rangle^t$

Extinction probability as a function of K

- Properties of $f_K(z)$ in [0,1]
 - $\begin{array}{l} \star \ f_K(0) = p_0 \\ \star \ f_K(1) = \\ \star \ f'_K(1) = \\ \star \ f'_K(z) \\ \star \ f''_K(z) \end{array}$

Extinction probability as a function of K



Phase transition in branching processes

• The fixed point condition for the probability of non-extinction $\rho = 1 - P_{extinction}$,

$$P_{extinction} = 1 - \rho = f_K(P_{extinction}) = f_K(1 - \rho) = \sum_{k=0}^{\infty} p_k(1 - \rho)^k$$

(because $P{A} + P{noA} = 1$). Expanding using the binomial theorem

$$1 - \rho = \sum_{k=0}^{\infty} p_k \left[1 - k\rho + \frac{1}{2}k(k-1)\rho^2 - \dots \right] = \blacksquare$$

$$= \sum_{k=0}^{\infty} p_k - \left(\sum_{k=0}^{\infty} p_k k\right) \rho + \frac{1}{2} \left(\sum_{k=0}^{\infty} p_k k(k-1)\right) \rho^2 + \ldots = \mathbb{I}$$
$$= 1 - \langle K \rangle \rho + \frac{1}{2} \langle K(K-1) \rangle \rho^2 + \ldots$$

• For small ρ (large $P_{extinction}$), introducing $\phi = \langle K(K-1) \rangle$ (2nd factorial moment)

$$\frac{1}{2}\phi\rho^2 - (\langle K \rangle - 1)\rho \simeq 0$$

which has 2 solutions,

$$ho = 0$$
 and $ho \simeq 2 rac{\langle K
angle - 1}{\phi}$

We need to consider the solution closer to (but smaller than) 1, so

$$\rho = 0 \quad \text{ for } \langle K \rangle \leq 1 \quad \text{ and } \quad \rho \simeq 2 \frac{\langle K \rangle - 1}{\sigma_c^2} \quad \text{ for } \langle K \rangle \geq 1$$

where we have used $\phi = \sigma^2 + \langle K \rangle (\langle K \rangle - 1)$, if $\rho \simeq 0$ then $\langle K \rangle \simeq 1$ and $\phi \simeq \sigma_c^2$

• The transition is continuous, but sharp \Rightarrow 2nd order phase transition The case $\langle K \rangle = 1$ is critical, as it separates two very different behaviors

Universality: close to the critical point



 $m = \langle K \rangle$

Continuous (or second order) phase transition

Let m be a control parameter (⟨K⟩ in branching or temperature, etc.)
 Let ρ be an order parameter (non-extinction probability, magnetization, etc.)
 Then

 $\rho \propto \begin{cases}
0 & \text{for } m \text{ below } m_c = \text{critical point} \\
(m - m_c)^{\beta} & \text{for } m \text{ above but close to } m_c
\end{cases}$

- Abrupt change in the derivative The derivative is discontinuous if $\beta \leq 1$
- For a branching process, $m_c = 1$ and $\beta = 1$



- For a magnetic system, m is the inverse of the temperature, ρ is magnetization
- \Rightarrow m_c is the inverse of Curie temperature and $\beta = 1/3$



Magnetization dissappears sharply

Heller & Benedek

Example: binomial number of offsprings

• Each element has only a fixed number of trials n to generate other elements

$$p_k = P\{K = k\} = \begin{pmatrix} n \\ k \end{pmatrix} p^k q^{n-k}, \text{ for } k = 0, 1, \dots n.$$

with p the probability of being successful in each trial, and $q\!=\!1-p$

• The probability generating function

$$f_K(z) = \sum_{k=0}^n \begin{pmatrix} n \\ k \end{pmatrix} q^{n-k} p^k z^k = (q+pz)^n.$$

using the binomial theorem. We will consider n=2

• $P_{extinction}$ will come from the smallest solution in [0,1] of

$$z^* = (q + pz^*)^2 \Rightarrow z^* = \frac{1 - 2pq \pm \sqrt{(1 - 2pq)^2 - 4p^2q^2}}{2p^2}$$

but for the square root we can write $\sqrt{1-4p(1-p)} = \sqrt{(1-2p)^2} = (1-2p)$

$$\Rightarrow z = \frac{1 - 2p + 2p^2 \pm (1 - 2p)}{2p^2} = \begin{cases} (1 - 2p + p^2)/p^2 &= (q/p)^2\\ p^2/p^2 &= 1 \end{cases}$$

The smallest root depends on whether p is below or above 1/2

$$P_{extinction} = \left\{ egin{array}{cc} 1 & \mbox{ for } p \leq 1/2 \ (q/p)^2 & \mbox{ for } p \geq 1/2 \end{array}
ight.$$

As $\langle K\rangle=np=2p$ the critical case $\langle K\rangle=1$ corresponds to $p=p_{C}=1/2$ (in agreement with the behavior of $P_{extinction})$

- In terms of the non-extinction probability $\rho = 1 - P_{extinction}$



Total size of the population

• The size of the population, summing across generations is

$$S = \sum_{t=0}^{\infty} N_t,$$

- \star total number of individuals that have ever existed, or
- total number of domino pieces toppling,
- ★ "size" of an earthquake, etc...
- Its mean value, for $\langle K \rangle < 1$, using the geometric series, and $\langle N_t \rangle = \langle K \rangle^t$ (new!)

$$\langle S \rangle = \langle N_0 \rangle + \langle N_1 \rangle + \langle N_2 \rangle + \ldots = 1 + \langle K \rangle + \langle K \rangle^2 + \ldots = \frac{1}{1 - \langle K \rangle}$$

Note that when $\langle K \rangle \rightarrow 1$, the probability of extinction is 1, but $\langle S \rangle \rightarrow \infty$ (!)

Total size of the population: binomial case

• Each element has only a fixed number of trials n to generate other elements

$$p_k = P\{K = k\} = \left(egin{array}{c} n \ k \end{array}
ight) p^k q^{n-k}, ext{ for } k = 0, 1, \dots n.$$

with p the probability of being successful in each trial, and $q\!=\!1-p$

- Remember $\langle K \rangle = np$, so the critical point is at $p_c = 1/n$
- Representation of a branching process as a tree (connected graph with no loops).
 - * Each element is associated to a node
 - * Branches linking nodes indicate an offspring relationship between two nodes



- Representation of a branching process as a tree (connected graph with no loops).
 - ★ Each element is associated to a node
 - * Branches linking nodes indicate an offspring relationship between two nodes
- All nodes have just one incoming branch, except the one in the zero generation
 - \star the number of branches is the number of nodes minus 1, i.e., s-1
 - \star the number of possible branches arising from s nodes is ns (in a n-tree)
 - \star the number of missing branches (non-successful trials) is ns (s 1)

A particular tree of size \boldsymbol{s} comes with a probability

$$p^{s-1}(1-p)^{(n-1)s+1}$$
 with $s=1,2,\ldots$
• For n = 2, the probability of having an undefined tree of size $s = 1, 2 \dots$ comes from the Catalan numbers! ...

$$P\{S=s\} = C_s p^{s-1} (1-p)^{s+1} = \frac{1}{s+1} \begin{pmatrix} 2s \\ s \end{pmatrix} p^{s-1} (1-p)^{s+1}$$

with $C_s = \frac{1}{s+1} \begin{pmatrix} 2s \\ s \end{pmatrix}$ the number

of different trees of size s, called Catalan numbers

The trees are the internal part of rooted binary trees

Can you draw them?



Calculation of the Catalan numbers

- Let us decompose a tree of size s into its root (zeroth generation) and the rest This can be done as
 - \star A subtree of size s-1 in the 1st branch and another of size 0 in the 2nd
 - \star A subtree of size s-2 in the 1st branch and another of size 1 in the 2nd
 - * ...
 - $\star\,$ A subtree of size 0 in the 1st branch and another of size s-1 in the 2nd
 - So, the total number of trees of size s is

$$C_s = C_0 C_{s-1} + C_1 C_{s-2} + \dots + C_{s-2} C_1 + C_{s-1} C_0 \qquad \text{with } C_0 = 1$$

• We define a generating function for the Catalan numbers

$$h(x) = C_0 + C_1 x + C_2 x^2 + \ldots = \sum_{s=0}^{\infty} C_s x^s$$

The properties of the Catalan numbers will allow the calculation of h(x)

$$[h(x)]^{2} = \sum_{i,j=0}^{\infty} C_{i}C_{j}x^{i+j} = \sum_{s=0}^{\infty} \underbrace{\left[\sum_{i+j=s}^{\infty} C_{i}C_{j}\right]}_{C_{s+1}} x^{s} = \frac{1}{x}\sum_{s=0}^{\infty} C_{s+1}x^{s+1} = \frac{h(x) - C_{0}}{x}$$

SO

$$h(x) = \frac{1 \pm \sqrt{1 - 4x}}{2x}$$

but this tell us nothing yet. Using the Taylor expansion of $\sqrt{1-x}$

$$\sqrt{1-x} = 1 - \frac{x}{2} - \frac{1}{4} \left(\frac{x^2}{2!} \right) - \frac{3}{8} \left(\frac{x^3}{3!} \right) - \dots = 1 - \frac{x}{2} - \sum_{s=1}^{\infty} \frac{(2s-1)!!}{2^{s+1}(s+1)!} x^{s+1}$$

then

$$\sqrt{1-4x} = 1 - 2x - \sum_{s=1}^{\infty} \frac{(2s-1)!!2^{s+1}}{(s+1)!} x^{s+1}$$

and so, taking the minus sign (otherwise h(x) is not a g.f.)

$$h(x) = \frac{1 - \sqrt{1 - 4x}}{2x} = 1 + \frac{1}{2x} \sum_{s=1}^{\infty} \frac{(2s - 1)!!2^{s+1}}{(s+1)!} x^{s+1} = 1 + \sum_{s=1}^{\infty} \frac{(2s - 1)!!2^s}{(s+1)!} x^s$$

then the Catalan numbers are, and using $(2s)! = (2s)!!(2s-1)!! = s!2^s(2s-1)!!$

$$C_s = \frac{(2s-1)!!2^s}{(s+1)!} = \frac{(2s)!}{s!(s+1)!} = \frac{1}{s+1} \begin{pmatrix} 2s \\ s \end{pmatrix}$$

the latter being valid for $s=0,1,2\ldots$

Coming back to the Taylor expansion of $\sqrt{1-4x}$

$$\sqrt{1-4x} = 1 - 2x \sum_{s=0}^{\infty} C_s x^s$$

Parenthesis: many uses of the Catalan numbers

Davis, geometer 2010

• Number of balanced configurations with n pairs of parenthesis

n = 0:	*	1 way
n = 1:	()	1 way
n = 2:	()(),(())	2 ways
n = 3:	()()(), ()()), (()(), (()()), ((()))	5 ways
n = 4:	()()(), ()(()), ()(()), ()(()), ()(())),	14 ways
	(())()(), (())(()), (()())(), ((()))(), (()())),	
	(()(())), ((())()), ((()())), (((())))	

• Number of mountains profiles with n upstrokes and n downstrokes

n = 0:	*	1 way
n = 1:	\land	1 way
n=2:	\wedge	2 ways
	/\/ / \	
n = 3:	\wedge	5 ways
	/\/\ /\/ / \/ / / \	

• Number of paths above (or on) the diagonal in a $n \times n$ lattice



• Number of triangulations of polygons with n+2 sides

• Non-crossing hand-shaking configurations of 2n people in a round table



• Many more!

Stanley Enumerative Combinatorics 1999

Normalization of the size distribution

• $P\{S=s\}$ is normalized for $p\leq 1/2$ but not for p>1/2

$$\sum_{s=1}^{\infty} P\{S=s\} = \frac{q}{p} \sum_{s=1}^{\infty} C_s p^s q^s = \frac{q}{p} [h(pq) - 1]$$

with q = 1 - p and introducing $h(x) = \sum_{s=0}^{\infty} C_s x^s$. As $h(x) = (1 - \sqrt{1 - 4x})/(2x)$

$$h(pq) = \frac{1 - \sqrt{1 - 4pq}}{2pq} = \frac{1 - \sqrt{(1 - 2p)^2}}{2pq} = \frac{1 - |1 - 2p|}{2pq} = \begin{cases} \frac{2p}{2pq} = \frac{1}{q} & \text{if } 1 \ge 2p \\ \frac{2(1 - p)}{2pq} = \frac{1}{p} & \text{if } 1 \le 2p \end{cases}$$

Therefore

$$\sum_{s=1}^{\infty} P\{S=s\} = \frac{q}{p}[h(pq)-1] = \begin{cases} \frac{q}{p}(\frac{1}{q}-1) = 1 & \text{if } p \le 1/2 \\ \frac{q}{p}(\frac{1}{p}-1) = (q/p)^2 & \text{if } p \ge 1/2 \end{cases}$$

which turns out into

$$\sum_{s=1}^{\infty} P\{S=s\} = P_{extinction}$$

But how does $P\{S = s\}$ look like? And what this has to do with power laws?

• Summarizing, the size distribution

$$P\{S=s\} = C_s p^{s-1} (1-p)^{s+1} = \frac{1}{s+1} \begin{pmatrix} 2s \\ s \end{pmatrix} p^{s-1} (1-p)^{s+1}$$

for a branching process with binomial distribution and $n\!=\!2$

• But what this has to do with power laws??

Asymptotic total size of the population

• Using Stirling's approximation, valid for $s \to \infty$ Christensen & Moloney 2005; A.C. & Font-Clos 2013

$$s! \sim \sqrt{2\pi s} \left(\frac{s}{e}\right)^s$$

the binomial coefficient turns out to be

$$\left(\begin{array}{c}2s\\s\end{array}\right) = \frac{(2s)!}{s!s!} \sim \frac{4\pi s}{2\pi s} \frac{(2s)^{2s}}{s^{2s}} \sim \frac{4^s}{\sqrt{\pi s}}$$

and the Catalan number, replacing $s+1 \sim s$

$$C_s = \frac{1}{s+1} \begin{pmatrix} 2s \\ s \end{pmatrix} \sim \frac{4^s}{\sqrt{\pi} \, s^{3/2}}$$

essentially, an exponential increasing function of \boldsymbol{s}

• Introducing the factor $p^{s-1}q^{s+1}$ we get $P\{S=s\}$

$$P\{S=s\}\sim \frac{q}{\sqrt{\pi}p}\frac{(4pq)^s}{s^{3/2}}$$

How does this function looks like for large s?

★ If $p(1-p) < 1/4 \Rightarrow p \neq 1/2 \Rightarrow$ decreasing exponential ★ If $p(1-p) = 1/4 \Rightarrow p = 1/2 \Rightarrow$ exponential dissapears \Rightarrow power law!

It becomes more transparent writting

$$(4pq)^s = e^{s \ln[4p(1-p)]} = e^{-s/\xi(p)}$$

with the characteristic size defined as

$$\xi(p) = \frac{-1}{\ln[4p(1-p)]} = \left(\ln\frac{1}{4p(1-p)}\right)^{-1}$$

and then

$$P(S=s) \sim \frac{q}{\sqrt{\pi}p} \frac{e^{-s/\xi(p)}}{s^{3/2}}$$

Case $p \neq 1/2$

★ For s large but s ≪ ξ(p) ⇒ power law with exponent 3/2
 ★ For s large with s ≫ ξ(p) ⇒ exponential decay

Case p = 1/2

 $\star\,$ Then, $\xi\!\rightarrow\!\infty$ and for large s we obtain a power law

The critical exponent for the size distribution is 3/2



Divergence of the characteristic size

• Another critical exponent arises for the divergence of $\xi(p)$ at the critical point. Introducing the deviation with respect to the critical point, $\Delta \equiv p - p_c = p - 1/2$

$$p(1-p) = \left(\frac{1}{2} + \Delta\right) \left(\frac{1}{2} - \Delta\right) = \frac{1}{4} - \Delta^2$$

So, close to the critical point (for small Δ)

$$\frac{1}{4p(1-p)} = \frac{1}{1-4\Delta^2} \simeq 1 + 4\Delta^2 + \dots$$

(using the formula of the geometric series), then

$$\ln \frac{1}{4p(1-p)} \simeq \ln(1+4\Delta^2) \simeq 4\Delta^2 + \dots$$

(using the Taylor expansion of the logarithm at point 1), therefore

$$\xi(p) = \left(\ln\frac{1}{4p(1-p)}\right)^{-1} \simeq \frac{1}{4\Delta^2} + \dots$$

So, $\xi(p)$ diverges at the critical point as a power law, with an exponent =2Then, for s large and Δ small

$$P(S=s) \sim \frac{1}{\sqrt{\pi}} \frac{e^{-4(p-p_c)^2 s}}{s^{3/2}}$$

Expected value of the size

• We already know that for $\langle K \rangle < 1$ (i.e., p < 1/2, i.e., $\Delta < 0$)

$$\langle S \rangle = \langle N_0 \rangle + \langle N_1 \rangle + \langle N_2 \rangle + \ldots = 1 + \langle K \rangle + \langle K \rangle^2 + \ldots = \frac{1}{1 - \langle K \rangle} = -\frac{1}{2\Delta}$$

substituting $\langle K \rangle = 2p$ and $\Delta = p - 1/2 =$ deviation with respect criticality This defines another critical exponent

- As $\xi(p)\,{\simeq}\,\Delta^{-2}/4$ close but below $p_c\,{=}\,1/2$ then

$$\xi(p) \simeq \langle S \rangle^2$$

So, if the mean increases by 2, the extreme values given by ξ increase by 4

Total size of the population: general case

• Let $g(z) = f_S(z)$ be the generating function of $S = \sum_{t=0}^{\infty} N_t$. Then

$$g(z) = z f_K(g(z))$$

with
$$f_K(z)$$
 the p.g.f. of the number of offsprings per element

- 1st demonstrated by Hawkins and Ulam in 1944 for nuclear chain reactions (as a part of the Manhattan project)
 - * A neutron may produce a fission reaction
 - $\star~$ Each reaction releases neutrons
 - ★ Each neutron may trigger more reactions, and so on.

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• Demonstration

Consider the size from generation 1 to ∞ (excluding the 0-th generation)

$$S_{\tilde{0}} = S - 1 = \sum_{t=1}^{\infty} N_t$$

with $q_s = P(S_{\widetilde{0}} = s)$ and a generating function $\tilde{g}(z) = \sum_{\forall s} q_s z^s$ A size s in generations from 1 to ∞ can be decomposed into

- $\star~$ a size k in the first generation, with probability p_k , and
- * a size s k in the remaining generations (from 2 to ∞) but starting with k elements; this has a probability $q_{s-k}^{(k)}$

(note that
$$q_{s} = q_{s}^{(1)}$$
)

Using the law of total probability,

$$q_s = \sum_{k=1}^s p_k q_{s-k}^{(k)},$$

except for s=0, where $q_0=p_0$

If we multiply by z^s and sum for all s we will obtain the g.f. of $S_{\widetilde{0}}$

$$\tilde{g}(z) = p_0 + \sum_{s=1}^{\infty} \sum_{k=1}^{s} p_k q_{s-k}^{(k)} z^s = p_0 + \sum_{k=1}^{\infty} p_k \left[\sum_{s=k}^{\infty} q_{s-k}^{(k)} z^{s-k} \right] z^k$$

Hawkins & Ulam 1944; A.C. & Font-Clos, arXiv 2012

The term in [] is the g.f. of the size from 1 to ∞ generations but starting with k elements $(N_1 = k)$

As these k parents are independent of each other

 \Rightarrow size will be the sum of k independent random variables each with g.f. $\tilde{g}(z)$

This yields $\left[\tilde{g}(z)
ight]^k$ as the corresponding generating function,

$$[\tilde{g}(z)]^k = \sum_{s-k=0}^{\infty} q_{s-k}^{(k)} z^{s-k},$$

Substituting into the equation above

$$\tilde{g}(z) = p_0 + \sum_{k=1}^{\infty} p_k [\tilde{g}(z)]^k z^k = f_K(z\tilde{g}(z))$$

As $S = 1 + S_{\tilde{0}}$ we need to add an independent variable with g.f. = z(as N_0 takes the value 1 with probability 1) then, the g.f. of the size from generation 0 to ∞ is the product $z\tilde{g}$, so

$$g(z) = z \tilde{g}(z) = z f_K(z \tilde{g}(z)) = z f_K(g(z))$$

- Binomial case. Substituting $f_K(z) = \left(q + pz\right)^2$ then

$$g(z) = z f_K(g(z)) = z (q + pg(z))^2 \qquad \Rightarrow \qquad g(z) = \frac{1 - 2pqz \pm \sqrt{1 - 4pqz}}{2p^2 z}$$

Using the Taylor expansion for the square root

$$\sqrt{1 - 4pqz} = 1 - 2pqz - \sum_{s=1}^{\infty} \frac{(2s - 1)!!2^{s+1}}{(s+1)!} (pqz)^{s+1}$$

where we do not need to compute the Catalan numbers $C_{\mathcal{S}}\text{, so, taking ``-''}$

$$g(z) = rac{q}{p} \sum_{s=1}^{\infty} C_s (pqz)^s$$

From the coefficients we recover the probability distribution we knew

$$P\{S=s\} = C_s p^{s-1} q^{s+1}$$

• Geometric case. "Success" probability p and q=1-p and values $k=0,1,2\ldots\infty$

$$p_k = P\{K = k\} = q^k p \qquad \Rightarrow \qquad f_K(z) = \sum_{k=0}^{\infty} p_k z^k = p \sum_{k=0}^{\infty} q^k z^k = \frac{p}{1 - qz}$$

using the geometric series. The generating function for the size is

$$g(z) = zf_K(g(z)) = \frac{pz}{1 - qg(z)} \Rightarrow g(z) = \frac{1 - \sqrt{1 - 4pqz}}{2q} = pz + \sum_{s=2}^{\infty} C_{s-1}q^{s-1}p^s z^s$$

where we have used the following, with C_i the $i\!-\!{\rm th}$ Catalan number

$$\sqrt{1-4x} = 1 - 2x - \sum_{i=1}^{\infty} 2C_i x^{i+1}$$

Therefore, the size distribution (without binary trees!)

$$P\{S=s\} = C_{s-1}q^{s-1}p^s \qquad \sim \frac{1}{4\sqrt{\pi}q} \frac{(4pq)^s}{s^{3/2}} \ \text{ for } s \to \infty$$

so we again obtain a critical exponent $\,=3/2$ (and also the others)

• C_s also counts number of (non-necessarily-binary) trees with s edges



Davis 2010

• Normalization of the size distribution in the geometric case

$$\sum_{s=1}^{\infty} P\{S=s\} = C_{s-1}q^{s-1}p^s = \begin{cases} 1 & \text{if } q \le 1/2 \\ p/q & \text{if } q \ge 1/2 \end{cases}$$

which corresponds to the probability of extinction in the geometric case Note that $\langle K\rangle=q/p$, so $p_C=q_C=1/2$

- Another offspring distribution
 - $\star~$ 0 offsprings with probability q=1-p
 - $\star~$ 2 offsprings with probability p

Then $f_K(z) = q + p z^2$. The generating function for the size is

$$g(z) = zf_K(g(z)) = z(q + pg(z)^2) \quad \Rightarrow \quad g(z) = \frac{1 \pm \sqrt{1 - 4pqz^2}}{2pz} = \sum_{i=0}^{\infty} C_i p^i q^{i+1} z^{2i+1} + \frac{1}{2pz} = \sum_{i=0}^{\infty} C_i p^i q^{i+1} z^$$

Therefore

$$P\{S=s\} = C_{\underline{s-1}} p^{\underline{s-1}} q^{\underline{s+1}} q^{\underline{s+1}} \qquad \text{for } s = 1, 3, 5 \dots$$

So, C_i counts the number of rooted binary trees of size s=2i+1 Asymptotically we do not scape from the exponent 3/2

$$P\{S=s\}\sim \sqrt{rac{2q}{\pi p}}rac{(4pq)^{s/2}}{s^{3/2}} \; ext{ for } s
ightarrow \infty$$

 C_i counts the number of rooted binary trees of size s=2i+1



Finite size effects in branching processes

• Let us consider a limitation in the number of generations: t = 0, 1, ... L (this plays the role of boundaries)

The probability of extinction, with $f(z) \equiv f_K(z)$, will be

$$P_{ext}(L) = f^L(0) \qquad < \qquad P_{\infty} = \lim_{t \to \infty} f^t(0)$$

- Consider a very large number of generations, \boldsymbol{n}
 - \Rightarrow $f^n(0)$ will be close to $f^{\infty}(0) = P_{\infty}$

Let us Taylor expand $f(f^n(0))$ around the fixed point P_{∞}

$$f^{n+1}(0) = f(f^n(0)) = P_{\infty} + f'(P_{\infty}) (f^n(0) - P_{\infty}) + \dots$$

• Taking up to 2nd-order terms and arranging, the inverse of the distance is¹

$$c_{n+1} \equiv \frac{1}{P_{\infty} - f^{n+1}(0)} = \frac{c_n}{M} + \frac{C}{M^2}$$

with $M = f'(P_{\infty})$ and $C = f''(P_{\infty})/2$. Iterating

$$c_{n+\ell} = \frac{c_n}{M^{\ell}} + \frac{C(1-M^{\ell})}{M^{\ell+1}(1-M)}$$

In the subcritical case, $P_{\infty} = 1$, then $M = \langle K \rangle$ and $2C = \sigma^2 + \langle K \rangle (\langle K \rangle - 1)$, so

$$c_{n+\ell} = \frac{c_n}{\langle K \rangle^{\ell}} + \frac{\sigma^2 (1 - \langle K \rangle^{\ell})}{2 \langle K \rangle^{\ell+1} (1 - \langle K \rangle)} - \frac{1 - \langle K \rangle^{\ell}}{2 \langle K \rangle^{\ell}}$$

¹do not confuse distance to the fixed point with distance to the critical point

• Let us introduce a rescaled distance to the critical point $y = \ell(\langle K \rangle - 1)$, so

$$c_{n+\ell} = \frac{\sigma^2 (1 - \langle K \rangle^{\ell})}{2 \langle K \rangle^{\ell+1} (1 - \langle K \rangle)} + \ldots = -\frac{\sigma_c^2 (1 - e^y)\ell}{2e^y y}$$

with $\langle K \rangle = 1 + y/\ell$ and $\langle K \rangle^{\ell} = e^{y}$ and with ℓ large (then $\langle K \rangle$ is close to 1) For $L = \ell + n \gg n$, we have that the probability of non-extinction will be

$$\rho(L) = 1 - P_{ext}(L) = 1 - f^{L}(0) = \frac{1}{c_{L}} = \frac{2e^{y}y}{\sigma_{c}^{2}(e^{y} - 1)L},$$

with $L \simeq \ell$. So, a scaling law is fulfilled, with scaling function $\mathcal{G}(y)$

$$\rho(L) = \frac{1}{L\sigma_c^2} \mathcal{G}(L(\langle K \rangle - 1)) \quad \text{ with } \quad \mathcal{G}(y) = \frac{2ye^y}{e^y - 1}$$

valid also for the supercritical case. This is called finite-size scaling

• Let us repeat



Phase transitions only exist in the infinite-system limit (thermodynamic limit)

$$\mathcal{G}(y) = \frac{2ye^y}{e^y - 1} \quad \rightarrow \quad$$

 $\begin{array}{ccc} 0 & \text{when } y \to -\infty \\ 2 & \text{when } y \to 0 \\ 2y & \text{when } y \to \infty \end{array}$

So, for $L\!\rightarrow\!\infty$

$$\rho(L) = \frac{1}{L\sigma_c^2} \mathcal{G}(L(\langle K \rangle - 1)) \quad \rightarrow$$

$$\left\{ \begin{array}{ll} 0 & \mbox{for } \langle K \rangle < 1 \\ 2\sigma_c^{-2}L^{-1} & \mbox{for } \langle K \rangle = 1 \\ 2(\langle K \rangle - 1)/\sigma_c^2 & \mbox{for } \langle K \rangle > 1 \end{array} \right.$$



Simulation of a branching process

- Initialize t = 0 and $N_0 = 1$ (one single ancestor)
- Loop for t
 - \star Simulate N_t values of K_i
 - \star Compute $N_{t+1} = \sum_{i=1}^{N_t} K_i$
 - \star If $N_{t+1} = 0 \Rightarrow$ stop
 - \star t = t + 1
- For the twins-or-nothing example

$$K = \begin{cases} 2 & \text{if } u \leq p \\ 0 & \text{otherwise} \end{cases}$$

with \boldsymbol{u} a uniform random number between 0 and 1

As the mean of the number of offsprings is $\langle K\rangle\,{=}\,2p$, then, $p_c\,{=}\,1/2$



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As the mean of the number of offsprings is $\langle K\rangle\,{=}\,2p$, then, $p_c\,{=}\,1/2$



Plot of $P\{S=s\}$ (with $\langle n \rangle = \langle K \rangle$)

As the mean of the number of offsprings is $\langle K\rangle\,{=}\,2p$, then, $p_c\,{=}\,1/2$



Plot of $P\{S=s\}$ (with $\langle n \rangle = \langle K \rangle$)

Earthquakes and branching processes

- Gutenberg-Richter power law holds only for ⟨K⟩ = 1
 Critical branching process ⇒ Fine tuning of mean number of offsprings
 ⇒ Very difficult to get in practice!
- Agreement only qualitative, not quantitative

$$1 + \frac{2b}{3} \simeq 1.67 \qquad \neq \qquad \frac{3}{2}$$



- Model too simple, still
- Kagan: Gutenberg-Richter exponent should be 3/2 (i.e., b-value = 0.75) Instrumental artifacts makes the exponent increase Kagan, Tectonophys 2010

Consequences for predictability

• Consider $\langle N_{t+1}|N_t \rangle$ with N_t known, then

 $\langle N_{t+1}|N_t\rangle = \langle K\rangle N_t$

using $N_{t+1} = \sum_{i=1}^{N_t} K_i$

For critical branching processes $\langle K\rangle = 1$ and then

 $\langle N_{t+1}|N_t\rangle = N_t$

Note that it is not only that the outcome of the next step is random It is much worst: the earthquake is in the limit of attenuation and intensification

• But what makes earthquakes critical?
Summary

- The size (energy) of earthquakes (and other natural hazards) follows a power-law distribution
- A power law signals the absence of a characteristic scale
- (Decreasing) power-law densities, with $\beta \leq 2$ have an infinite mean value
- Galton-Watson branching process can be a model of earthquakes
 - **\star** Continuous phase transition at $\langle K \rangle = 1$
 - Size distribution is only power law at the critical point





Self-Organized Branching Process

• Consider: 0 offsprings with prob 1-p 2 offsprings with prob p

Limit the maximum number of generations \Rightarrow analogous to introduce a boundary at t = L

Change p from one realization T to the next as

$$p(T+1) = p(T) + \frac{1 - N_L(T)}{M}$$



where N_L is the population in the last generation (=2 in Fig.) and M is a big number (explained later)

Note that there are 2 times scales

$$\star$$
 $t =$ fast time scale, counts generations , from $t = 0$ to L
 \star $T =$ slow time scale, counts realizations

$$p(T+1) = p(T) + \frac{1 - N_L(T)}{M}$$

Dynamics

- ★ If p is low \Rightarrow small size \Rightarrow $N_L = 0 \Rightarrow p$ increases ★ If p is high \Rightarrow large size \Rightarrow $N_L > 1 \Rightarrow p$ decreases
- Indeed, we know that $\langle N_L \rangle = \langle K \rangle^L = (2p)^L$ So, we can write, $N_L = (2p)^L + \eta$, with $\langle \eta \rangle = 0$ Considering the deterministic equation (removing η)

$$p(T+1) = F(p(T)) = p(T) + \frac{1 - (2p(T))^L}{M}$$

Therefore, the deterministic equation has a fixed point $p^* = 1/2 = p_c$

$$1 - p$$

• Moreover, if M is big enough then $|F'(p^*)| < 1$ and the fixed point is attractive, so

 $p(T) \to p^* = p_c$

As the noise is small, it only adds small perturbations to p^{\ast}

Then, p tends, or self-organizes, to its critical value, on average

• Note:

Self-organization is the spontaneous emergence of structures or global order (here we do not have any structure yet, but wait...)

Examples:

convection patterns in fluids, chemical oscillations, self-regulations of markets

• Nevertheless, the global condition (on p) is very difficult to justify, in practice

Cellular automaton Manna model²

- Let us consider a lattice in d dimensions
 - ★ Each site can store only 1 particle (or 0)
 - \star If extra particles arrive at a site:



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Bak 1996, after Grassberger
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- ⇒ 2 of them are transferred to 2 randomly chosen sites among its neighbors (this may generate an avalanche)
- ★ Particles leave the system through the (open) boundary
- \star If there is no activity (all sites with 1 particle or less):
 - \Rightarrow Add 1 particle to a random site

In a formula, with nn(j) denoting 2 random neighbors of j

if
$$z_j \ge 2 \qquad \Rightarrow \begin{cases} z_j & \to & z_j & - & 2 \\ z_{nn(j)} & \to & z_{nn(j)} & + & 1 \end{cases}$$

 $\text{if } z_k < 2 \ \forall k \quad \Rightarrow \qquad z_n \rightarrow z_n + 1 \ \text{ with } n = rand \\$

²Cellular automaton = dynamical system with discrete time, space, and variable (field)

• The Manna model defines a complex system:

System composed of many interacting parts, such that the collective behavior of those parts together is more than the sum of their individual behaviors

Other examples, more complex: the cell, the brain, ecosystems, the economy, the Earth's crust...

• Let us go back to the Manna model in the limit of infinite dimensions, $d
ightarrow \infty$

Then, the propagation of the activity will show no loops \simeq mean field (a neighbor will not be selected twice to get a grain \Rightarrow no overlap)

So, there will be no spatial correlations, and all sites are equivalent (the boundary conditions need to be readjusted)

Each site will become active $(z \ge 2)$ with the same probability

p =fraction of sites with one particle = P[z = 1]

- Then, the activity propagates through the system as a branching process
- The offspring distribution will be binomial, with n=2 and parameter pBut note that there is no pre-existing tree
- The total number of particles will evolve as

$$\mathsf{mass}(T+1) = \mathsf{mass}(T) + 1 - \mathsf{out}(T)$$

(one particle added before the avalanche, "out" particles lost at the boundaries) Dividing by the total number of sites M, with p = P[z = 1] = mass/M

$$p(T+1) = p(T) + \frac{1 - \operatorname{out}(T)}{M}$$

which corresponds to the self-organized branching process Christensen & Moloney 2005 The evolution and adjustment of p is implemented in a natural way

Self-Organized Criticality (SOC) Bak et al. Phys Rev Lett 1987

• The dynamics arises from the sandpile metaphor

★ If there are few grains (flat pile)
⇒ small avalanches, pile grows
★ If there are many grains (steep pile)
⇒ large avalanches, pile decreases (through boundary dissipation)

This mechanism makes the slope of the pile fluctuate around the critical state

⇒ Bak-Tang-Wiesenfeld (BTW) model





• BTW model: one-dimensional lattice, d = 1, with $j = 1 \dots L$ Modification: no random selection of neighbors

$$\begin{array}{lll} \text{if } z_j \geq 2 & \Rightarrow & \left\{ \begin{array}{ll} z_j & \to & z_j & - & 2 & \text{for } j \neq L \\ z_{j\pm 1} & \to & z_{j\pm 1} & + & 1 \end{array} \right. \\ \text{if } z_k < 2 \ \forall k & \Rightarrow & z_n \rightarrow z_n + 1 \ \text{ with } n = rand \end{array}$$

The "particles" are in fact elements of slope in a 2-d sandpile

height at $j = h_j = h_{j+1} + z_j \implies z_j = h_j - h_{j+1}$

with $h_{L+1} = 0 \Rightarrow z_L = h_L \Rightarrow z_L \to z_L - 1$ (conserved BC)

 $\begin{array}{ll} \text{if } h_j - h_{j+1} \ge 2 & \Rightarrow & \left\{ \begin{array}{ll} h_j & \to & h_j & - & 1 \\ h_{j+1} & \to & h_{j+1} & + & 1 \end{array} \right. \\ \text{if } h_k - h_{k+1} < 2 \ \forall k & \Rightarrow & h_m \to h_m + 1 \ \text{for } m \le n \ \text{with } n = rand \end{array}$



• Height *h* picture (grains) versus slope *z* picture (repelling particles)



Christensen & Moloney 2005

Relation with interface depinning

Paczuski & Boettcher, Phys Rev Lett 1996

Define H_j as the total number of topplins in a sandpile When:

- * the initial condition is empty $(h_j = 0 \text{ for all } j \text{ and for } T = 0)$ and
- \star the addition takes place at j = 1

then, H defines an advancing interface, whose gradient gives the pile height

$$h_j = H_{j-1} - H_j$$

with H_0 giving the total number of grains added

Retrospective of models

- Domino (Otsuka) model of fracture
- Galton-Watson branching process
- Self-organized branching model
- Cellular automaton Manna (bureaucrats) model
- Bak et al. sandpile model
- Interface depinning model
- These models serve as metaphors for earthquakes







• Inspiration: Critical Points of Thermodynamic Phase Transitions

Magnetic material: atom = spin with 2 states

There exists a critical temperature T_c

- \star Above T_c : no magnetization, small clusters
- \star Below T_c : magnetization, one very large cluster
- * At the precise value $T = T_c \Rightarrow$ clusters of all sizes \Rightarrow power law!



Christensen & Moloney, Complexity and Criticality 2005

Burridge-Knopoff spring-block models

Bull Seism Soc Am 1967

- Earthquakes take place in "pre-existing" faults
 - ⇒ Alternative: modeling friction in a fault
 - ★ Experiment: spring-block system pulled from one end



Computer simulations:
 All blocks connected by flat springs to a moving plate

- stick-slip dynamics: slow driving (pull) + fast avalanches (shocks)
 - * The force on the block(s) increases (linearly) very slowly
 - * At some time (for some block) the force exceeds the static frictional force
 - Then, that block moves fast, changing the force over the neighbor blocks and so on

"Size" of the earthquake $\,\simeq\,$ number of sliding blocks

Coupled-map lattice model



• Olami-Feder-Christensen (OFC) model,

 $\it Phys \ Rev \ Lett \ 1992$

Two-dimensional version of Burridge-Knopoff model

- \star Coil (helical) springs connecting blocks in the direction of motion of the plate
- Flat (leaf) springs connecting blocks in the perpendicular direction (making the force then in the direction of motion also)
- \star In both cases the value of the elastic constants is K
- ★ Flat springs connecting blocks with the upper moving plate with constants $K_L \neq K$

- Let us define
 - \star $F_{i,j}$ = Force on block i, j
 - ★ $x_{i,j}$ = Displacement in the direction of motion of i, j relative to the upper flat spring

Also, the zero force between each pair corresponds to the lattice of upper springs. By Hooke's law $F_{i,j} =$



$$= -K(x_{i,j} - x_{i-1,j}) - K(x_{i,j} - x_{i+1,j}) - K(x_{i,j} - x_{i,j-1}) - K(x_{i,j} - x_{i,j+1}) - K_L x_{i,j}$$

$$= K\left(\sum_{nn(i,j)} x_{nn(i,j)} - 4x_{i,j}\right) - K_L x_{i,j}$$

If the upper plate moves with constant (small) velocity v then

$$\frac{dF_{k,l}}{dt} = -K_L \frac{dx_{k,l}}{dt} = K_L v \quad \text{for all } k, l$$

• When the force on some block i, j reaches the frictional threshold force F_{th} \Rightarrow block i, j slips instantaneously to the position with of zero force, so

 $F_{i,j} \rightarrow 0$ (assumption of the model)

Then, if we denote the new position of i, j as $x'_{i,j}$

$$0 = K\left(\sum_{nn(i,j)} x_{nn(i,j)} - 4x'_{i,j}\right) - K_L x'_{i,j}$$

where nn(i, j) denotes the nearest neighbors of i, j. Substracting,

$$F_{i,j} - 0 = -(4K + K_L)(x_{i,j} - x'_{i,j})$$

• Therefore, the force on the i + 1, j neighbor (for instance)

$$F_{i+1,j} = K\left(\sum_{nn(i+1,j)} x_{nn(i+1,j)} - 4x_{i+1,j}\right) - K_L x_{i+1,j}$$

So, as
$$F_{i,j} = -(4K + K_L)(x_{i,j} - x'_{i,j})$$
 then $F_{i+1,j}$ changes to

$$F_{i+1,j} \to F_{i+1,j} + K(x'_{i,j} - x_{i,j}) = F_{i+1,j} + \frac{K}{4K + K_L}F_{i,j}$$

and the model is non-conservative, as $\alpha = K/(4K + K_L) < 0.25$ except if $K_L \rightarrow 0$

Summary of the rules of the OFC model

if $F_{i,j} < F_{th}$ for all $i, j \Rightarrow dF_{i,j}/dt = K_L v$ with v very small

if
$$F_{i,j} \ge F_{th}$$
 for some $i, j \Rightarrow \begin{cases} F_{nn(i,j)} \to F_{nn(i,j)} + \alpha F_{i,j} \\ F_{i,j} \to 0 \end{cases}$

The boundary conditions are disregarded

Note that there are 2 times scales:

The slow one is continuous, but the fast one is discrete

In practice, in simulations, don't use $dF_{i,j}/dt = K_L v$. Why?

Then, the slow time scale turns into discontinuous

 \Rightarrow coupled map lattice model³

³Coupled map lattice = dynamical system on a lattice with continuous variables and discrete time

Earthquakes can be a SOC phenomenon

- Ingredients for SOC (and fulfillment in earthquakes)
 - \star Time scale separation (\Rightarrow OK)
 - ⋆ Thresholds, interaction
 - ★ Avalanche dynamics
 - Power-law distributions (with finite-size scaling)
 - ★ Underlying 2nd-order phase transition, reached by self-organization $(\Rightarrow ??)$

Think in the critical temperature of Fe, $T_c = 770^{\circ}$ C or in the critical point of water, at $T_c = 374^{\circ}$ C and 218 atm

Andrews 1869 (for CO₂)



Pruessner, private comm.

Other candidates for SOC

• For rain, Peters and Neelin have shown:

Nature Phys 2006, Neelin et al. Phil Trans R Soc A 2008

1. Existence in the atmosphere of a non-equilibrium stability-instability transition



Stanley, Rev Mod Phys 1999

• Finite size effects

Finite size scaling:
$$\langle P \rangle = L^{-0.2/\nu} H[(w - w_c)L^{1/\nu}]$$
 (L system size)
 $H(x) \propto \begin{cases} |x|^{-\gamma} & \text{for } x \to -\infty \\ x^{0.2} & \text{for } x \to +\infty \end{cases} \xrightarrow{\Rightarrow} \langle P \rangle \propto \begin{cases} 0 & \text{for } w < w_c \\ (w - w_c)^{0.2} & \text{for } w > w_c \end{cases}$

With critical point $w_c \simeq 63 \text{ mm}$ if T = 271 K, and so on

Phase transitions (abrupt changes) only exist in the limit $L \rightarrow \infty$



• Peters and Neelin have also shown:

Nature Phys 2006, Neelin et al. Phil Trans R Soc A 2008

2. The atmosphere is attracted towards the critical point of the transition



Continuous, non-upper-truncated power-law distributions

• Given by a probability density D(x) with x real (continuous)

$$D(x) = \frac{B}{x^{\gamma}}$$
 for $a \le x < \infty$

with a > 0. Then, $\int_a^\infty D(x) dx = 1$ with $\gamma > 1$ implies

$$B = (\gamma - 1)a^{\gamma - 1}$$

• In order to decide between competing explanations, universality classes, etc., it is important not only to determine if power laws hold, but also the precise value of the exponent γ

Fitting power-law distributions

- Some authors have pointed out the superiority of maximum-likelihood (ML) estimation

Goldstein et al. *Eur Phys J B*Bauke, *Eur Phys J B*White et al. *Ecol*Clauset et al. *SIAM Rev*

• ML estimators are: asymptotically unbiased and with lowest variance

Invariant under re-parameterizations



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Maximum likelihood (ML) estimation

• Given a dataset of size N, $x_1, x_2, \ldots x_N$, the likelihood is the joint distribution

$$\mathcal{L}(\gamma) = D(x_1, x_2, \dots, x_N; \gamma) = \prod_{i=1}^N D(x_i; \gamma)$$

(assuming independence). For a power law, $D(x) = B/x^{\gamma}$, the log-likelihood is

$$\ell(\gamma) = \frac{\ln \mathcal{L}(\gamma)}{N} = \frac{1}{N} \sum_{i=1}^{N} \ln D(x_i; \gamma) = \ln B - \frac{\gamma}{N} \sum_{i=1}^{N} \ln x_i = \ln B - \gamma \ln G$$

with G the geometric mean of the data and $B = (\gamma - 1)a^{\gamma - 1}$ (continuous case). Maximizing

$$\frac{d\ell}{d\gamma} = 0 \qquad \Rightarrow \qquad \gamma = 1 + \frac{1}{\ln \frac{G}{a}}$$

• Note that the ML solution depends only on the geometric mean ${\cal G}$

$$\gamma = 1 + \left(\ln\frac{G}{a}\right)^{-1}$$

So, any data, from any distribution, with the same G yields the same γ Then, maximum likelihood should be called minimum unlikelihood

⇒ A goodness-of-fit test is necessary

Goodness-of-fit test

• In order to test the goodness of the (ML) fit let us consider Kolmogorov-Smirnov

KS distance or KS statistic = maximum difference for all x between empirical S(x) and theoretical S(x)

$$d = \max\{S_{emp}(x) - S_{th}(x)\}$$

with $S(x) = \int_x^\infty D(x) dx$

Care with p-value: Monte Carlo simulations

• The problem of power-law fitting is not in fitting the power law





Clauset's et al. recipe

Clauset, Shalizi & Newman SIAM Rev 2009

- \star Take an arbitrary value of a (= minimum x for which the power law holds)
 - \star Calculate fit by ML estimation \Rightarrow yields exponent γ
 - Calculate Kolmogorov-Smirnov distance between empirical distribution and fit (no goodness-of-fit yet)
- \star Select the value of a which minimizes Kolmogorov-Smirnov distance $d = d_{emp}$ So, we come out with a fit given by a_{emp} and γ_{emp}
- \star Calculate p-value of the fit by simulating N_{sim} synthetic samples:
 - * Simulating a power law with exponent γ_{emp} for $x \ge a_{emp}$
 - * Bootstrap of the empirical distribution for $x < a_{emp}$
 - \star Proceed with synthetic samples in exactly the same way as with the empirical
- \Rightarrow Each synthetic sample yields a value of d_{sim}
- \Rightarrow Calculate p-value as $p \simeq \{$ number of $d_{sim} > d_{emp} \} / N_{sim}$

• Justification of the minimization of d

Under the null hypothesis, Kolmogorov-Smirnov distance goes as

$$d \propto rac{1}{\sqrt{N}}$$

1

So, under the null hypothesis, the smaller a, the larger N and the smaller dBut as soon as the null hypothesis fails, the fit deviates and d increases

A sort of balance between the two effects is implicit

Nevertheless, there is no reason why this deviation should compensate and overcome the reduction in \boldsymbol{d}

(it would depend on the shape of the distribution for x < a)

Problems of Clauset et al.'s recipe

• The method performs bad when generalized to truncated power laws

$$D(x) \propto \frac{1}{x^{\gamma}}$$
 for $a \leq x \leq b$

This is common in complex systems, due to finite-size effects



More problems of Clauset et al.'s recipe

A.C., Font & Camacho Phys Rev E 2011

• Consider nuclear half-lives: from below 10^{-16} s to 10^{23} yr $\sim 10^{31}$ s for 128 Te



(p should be uniformly distributed between 0 and 1 under the null hypothesis)



 \Rightarrow Failure of the Clauset's et al. recipe

Alternative recipe

Peters et al. J Stat Mech 2010; Deluca & A.C. Acta Geophys 2013

- \star Take an arbitrary value of a (= minimum x for which the power law holds)
 - \star Calculate fit by ML estimation \Rightarrow yields exponent γ
 - ★ Calculate KS distance d between empirical distribution and fit (no difference with Clauset et al. yet)
- \star Calculate a p-value for fixed a
 - * Simulate N_{sim} power-law synthetic samples with γ for $x \ge a$
 - \star Proceed with synthetic samples in exactly the same way as with the empirical
 - \Rightarrow Each synthetic sample yields a value of d_{sim}
 - \star Calculate p as

$$p \simeq rac{\operatorname{number of } d_{sim} > d}{N_{sim}}$$

• Select the smallest value of a provided that p > 0.20 (e.g.)

Performance

• Consider again nuclear half-lives: d and p(=q) versus a



We obtain $a_{emp} = 3 \times 10^7$ s (~ 1 yr) and $\gamma_{emp} = 1.09$ (with p > 0.20)
• Comparison between Clauset et al.'s solution (red) (power-law rejected) and alternative (green) (for the (complementary) cumulative distribution)



• Global earthquakes revisited: power law cannot be rejected



But other fits are possible! \Rightarrow LRT, or AIC, or BIC...



(Upper) Truncated power laws

• Deviations from a power law arise for large x, due to finite size effects. So,

$$D(x) = \frac{B}{x^{\gamma}} \quad \text{ with } a \leq x \leq b \quad \text{ and } B = \frac{(\gamma - 1)a^{\gamma - 1}}{1 - (a/b)^{\gamma - 1}}$$

Be careful: b is not b-value. The log-likelihood is now

$$\ell(\gamma) = \ln B - \gamma \ln G = -\gamma \ln \frac{G}{a} - \ln a + \ln \frac{\gamma - 1}{1 - (a/b)^{\gamma - 1}}$$

with G the geometric mean of data between a and b.

The log-likelihood needs to be maximized numerically

But the rest of the method is the same, swapping both a and b

Tropical cyclones

• (hurricanes, typhoons)



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$$\mathsf{Energy} \simeq \int \int C_D \rho |v(r,t)|^3 d^2 r dt$$

Bister & Emanuel, Met Atm Phys 1998



• Typhoons in the North Western Pacific (only the largest ones)



• Tropical cyclones (hurricanes, typhoons),

A.C., Ossó, Llebot, Nature Phys 2010



Peters et al. Stat Mech 2010



Discrete power laws

A. C., A. Deluca, & R. Ferrer-i-Cancho, ArXiv 2012

• The probability function is given by

$$f(n) = rac{B}{x^{\gamma}}$$
 with $x = a, a + 1, \dots$, and $\gamma > 1$ and $B = rac{1}{\zeta(\gamma, a)}$

where $\zeta(\gamma, a)$ is the Hurwitz theta function ($\zeta(\gamma, 1)$ is the Riemann function) The log-likelihood is

$$\ell(\gamma) = -\ln\zeta(\gamma, a) - \gamma\ln G$$

which is more difficult to maximize

Care with the cumulative distribution function (for the KS test)

The simulation of the discrete distribution is more involving also

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